Nanomineralogy of the First Solids in the Solar System: Discovering New Minerals and New Materials

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Nanomineralogy is the study of Earth and planetary materials at nano-scales, focused on characterizing nanofeatures (like inclusions, exsolution, zonation, coatings, pores) in minerals and revealing nanominerals and nanoparticles [1,2]. With advanced high-resolution analytical scanning electron microscopy, we are now capable to characterize solid materials in situ easier and faster down to nano-scales [3]. During our ongoing nanomineralogy investigation of meteorites since 2007, more than 20 new minerals have been discovered. Each of the new extraterrestrial minerals reveals distinctive forming environments with intensive variables (e.g., composition, temperature, pressure, fO₂), providing new insights into nebula or parent-body processes. 10 of them are refractory minerals, including allendeite (Sc₄Zr₃O₁₂) and hexamolybdenum (Mo,Ru,Fe) [4], tistarite (Ti₂O₃) [5], panguite [(Ti⁴⁺,Al,Sc,Mg,Zr,Ca, \Box)₂O₃] [6], kangite [(Sc,Ti,Al,Zr,Mg,Ca,n)₂O₃] [7], davisite (CaScAlSiO₆) [8], grossmanite (CaTi³⁺AlSiO₆) [9], paqueite [Ca₃TiSi₂(Al₂Ti)O₁₄] [10], and krotite (CaAl₂O₄) [11]. They are among the first solids formed in the solar nebula. To date, ~ 45 refractory minerals plus about 15 presolar minerals mark the beginning of the solar mineral evolution. There are now about 4800 mineral species identified on Earth. Minerals in the solar system evolved as a consequence of physical, chemical and biological processes over the past 4.568 billion years [12].

Presented here are a few new refractory minerals (including panguite, kangite, allendeite and tistarite) from the Allende meteorite and demonstrate how nanomineralogy works with an integrated EPMA-SEM-EDS-EBSD approach. EPMA, pioneered by Castaing, is one of the most important research tools in Earth sciences [13], which is responsible for discovery and/or chemical composition determination of new minerals since 1960s. Field emission SEM is widely available nowadays for nanoimaging with EDS for fast elemental analysis and EBSD for crystal structure and orientation studies. EBSD has been used successfully for structure determination of new minerals since 2006 [14].

Panguite (IMA 2010-057) is a high-temperature titania mineral with a cation-deficient *Pbca* bixbyite-related structure and a formula unit $(Ti,Al,Sc,Mg,Zr,Ca)_{1.8}O_3$, occurring with davisite (a newly-approved Sc-rich clinopyroxene [8]) in ultra-refractory inclusions from Allende [6] and Murchison, as shown in Fig. 1. The empirical chemical formula of type panguite by EPMA is $[(Ti_{0.75}Zr_{0.15}Si_{0.07})^{4+}_{\Sigma 0.97}(Al_{0.21}Sc_{0.20}Y_{0.06}V_{0.02}Cr_{0.01})^{3+}_{\Sigma 0.5}(Mg_{0.17}Ca_{0.10}Fe_{0.04})^{2+}_{\Sigma 0.31}]_{\Sigma 1.78}O_3$. EBSD patterns of panguite imply a CaF₂-type structure, which assisted in identifying its bixbyite-related structure (a defect form of the CaF₂ structure) by synchrotron diffraction [6].

Nature shows many advantages over technology on materials synthesis. The knowledge of minerals can be used for getting inspired on making new materials with useful properties [15]. Panguite and kangite are not only new minerals but also new materials from the solar nebula, which can be exploited for finding engineering materials.

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FIG. 1. Back-scatter electron image showing panguite with davisite in an ultra-refractory inclusion within an Allende amoeboid olivine aggregate.