

Maser Mapping of Red Supergiants and the Onset of Bipolar Outflow

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Maser emission from SiO, H₂O and OH is found in oxygen-rich red giant circumstellar envelopes at successively larger distances from the star. H₂O emission at 22 GHz has been mapped with MERLIN (the Multi Element Radio Linked Interferometer, <http://www.jb.man.ac.uk/merlin/MERLIN.html>) with a beamsize of 8 mas. Typically the 22 GHz emitting region is seen to be a thick expanding shell 100 mas across. Small-scale clumpiness is revealed. The shell is resolved into discrete features of velocity extent 1 to 3 km s⁻¹ and spatial extent 2 to 20 mas. Using this 3-D datacube, constraints are placed on the physical conditions such as kinetic temperature, density and local radiation field. These observations show that the size of the region increases with stellar mass-loss rate and that the mass outflow is undergoing acceleration. The outflow is gravitationally bound to the star at the inner radius of the H₂O masing region, but unbound at its outer radius (Chapman & Cohen 1986, *MNRAS*, 230, 415; Yates & Cohen 1994, *MNRAS*, 270, 529).

In NML Cyg (mass $\sim 50 M_{\odot}$), proper motion has been detected in two isolated features outside the rest of the H₂O maser shell which are symmetric in position and velocity with respect to the stellar position (Richards, Yates & Cohen 1996, *MNRAS*, 282, 665). The increase in separation over 9 years corresponds to a transverse velocity of 19 ± 6 km s⁻¹ at a distance of 2 kpc. This is consistent with the observed radial-velocity range of the 22 GHz emission. OH and dust observations show a similar axis of symmetry, and this may be a sign of the onset of bipolar outflow. Axisymmetric structures are often seen in supernova remnants, and also in the planetary nebulae which originate from $\sim 1 M_{\odot}$ stars. Results are also given for other supergiant stars currently being observed.

Papers relating to this work will be available by ftp; those interested should write to amsr@jb.man.ac.uk for details.