

The Far-Infrared Properties of the Most Isolated Galaxies

U. Lisenfeld^{1,2}, L. Verdes-Montenegro², J. Sulentic³, S. Leon⁴,
D. Espada^{2,5}, G. Bergond², E. García², J. Sabater²,
J. D. Santander-Vela², and S. Verley^{6,7,8}

¹Dept. Física Teórica y del Cosmos, Universidad de Granada, Spain
email: ute@ugr.es

²Instituto de Astrofísica de Andalucía (IAA/CSIC), Apdo. 3004, 18080 Granada, Spain

³Department of Astronomy, University of Alabama, Tuscaloosa, USA

⁴Instituto de Radioastronomía Milimétrica (IRAM), Avda. Divina Pastora 7, local 20, 18012 Granada, Spain

⁵Institute of Astronomy and Astrophysics, Academia Sinica, P.O. Box 23-141, Taipei 106, Taiwan

⁶GEPI/CAI, Observatoire de Paris, 77 avenue Denfert-Rochereau, 75014 Paris, France

⁷LERMA – Observatoire de Paris, 61 avenue de l’Observatoire, 75014 Paris, France

⁸INAF-Osservatorio Astrofisico di Arcetri, Largo E. Fermi 5, 50125 Firenze, Italy

Abstract. We describe the mid- (MIR) and far- (FIR) infrared properties of a large (~ 1000) sample of the most isolated galaxies in the local Universe. This sample is intended as a “nurture-free” zero point against which more environmentally influenced samples can be compared. We reprocess IRAS MIR/FIR survey data using the ADDSCAN/SCANPI utility for 1030 out of 1050 galaxies from the Catalogue of Isolated Galaxies (CIG) as part of the AMIGA project. We focus on diagnostics (FIR luminosity L_{FIR} , $R = \log(L_{\text{FIR}}/L_B)$ and IRAS colours) thought to be sensitive to effects of environment or interaction. The distribution of $\log(L_{\text{FIR}})$ is sharply peaked from 9.0–10.5 with very few ($< 2\%$) galaxies above 10.5. Review of available optical images of the most FIR luminous galaxies finds the majority to be, likely, interacting systems missed in our earlier morphological reevaluation. The optically normalised luminosity diagnostic $R = \log(L_{\text{FIR}}/L_B)$ shows a distribution sharply peaked between 0.0 and -1.0 . These results were compared to the magnitude limited CfA sample that was selected without environmental discrimination. This modestly (e.g. compared to cluster, binary galaxy and compact group samples) environmentally affected sample shows significantly higher mean $\log(L_{\text{FIR}})$ and R , whereas the mean $\log(L_B)$ is the same. Our sample shows a strong L_{FIR} vs. L_B correlation, with a slope steeper than one ($L_{\text{FIR}} \propto L_B^{1.41}$). Interacting galaxies were found above this correlation, showing an enhancement in L_{FIR} . With respect to the IRAS colours, we found higher F_{60}/F_{100} value for ellipticals and late-type galaxies than for spirals, indicating a higher dust temperature. The mean value of F_{60}/F_{100} was found to be lower than for interacting samples from the literature. The results indicate that the FIR emission is a variable enhanced by interaction, and that our sample probably shows the lowest possible mean value. This attests to the utility of our sample for defining a nurture-free zero point. More details can be found in the paper (*The AMIGA sample of isolated galaxies, III. IRAS data and infrared diagnostics*, U. Lisenfeld *et al.*) which is accepted for publication in *Astronomy & Astrophysics*.

Keywords. galaxies: evolution – galaxies: interactions – galaxies: luminosity function, mass function – galaxies: ISM – surveys – infrared: galaxies