

The Green Peas: Searching for LyC Emitters at Low Redshift

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Abstract. The escape fraction of Lyman continuum (LyC) radiation from galaxies remains one of the primary uncertainties in studies of reionization. However, few LyC-emitting galaxies are known. The recently identified, low-redshift “Green Pea” (GP) galaxies exhibit a number of similarities with high-redshift galaxies, and their optical emission lines suggest they may be some of the elusive LyC emitters. Recent *HST* COS and ACS observations of four GPs suggest further evidence for LyC escape and give new insights into the origins of Ly α and low-ionization UV lines in high-redshift galaxies. The Ly α emission and low-ionization emission and absorption lines provide a coherent physical picture of the neutral gas distribution in the GPs and may identify LyC emitters at high redshift. The rare, low-redshift GPs hint at possible factors that may enable LyC and Ly α escape from high-redshift galaxies.

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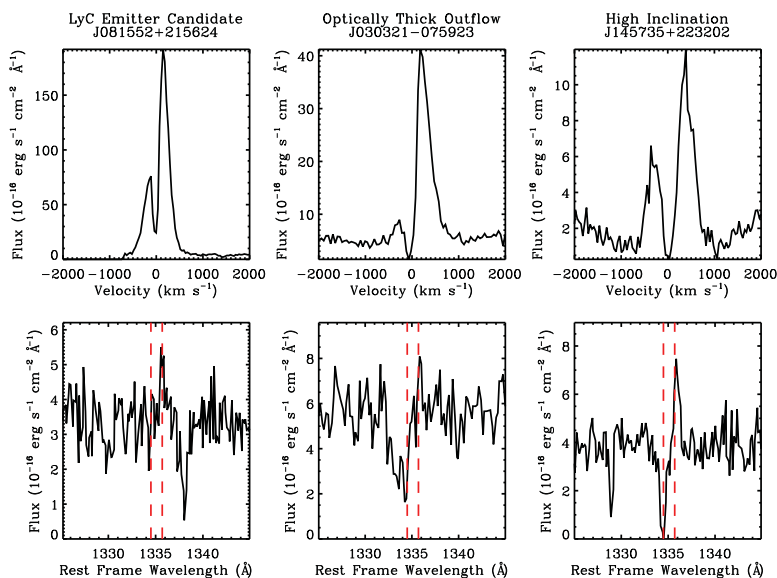


Figure 1. Top: Ly α emission from three Green Peas with different inferred optical depths and geometries. **Bottom:** The red lines show the expected positions of [C II] λ 1334.5 absorption and [C II]* λ 1335.7 emission. The LyC emitter candidates have strong, narrow Ly α , indicative of minimal scattering in neutral gas. They show scattered [C II]* emission but no line-of-sight [C II] absorption. The galaxy with an inferred optically thick outflow has deep, blue-shifted absorption and weaker, redshifted Ly α emission. The final Green Pea shows weak Ly α within an absorption trough and strong low-ionization absorption and emission from a high column density ISM.