LAST ICE AGE MILLENNIAL SCALE CLIMATE CHANGES RECORDED IN HUON PENINSULA CORALS

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ABSTRACT. Uranium series and radiocarbon ages were measured in corals from the uplifted coral terraces of Huon Peninsula (HP), Papua New Guinea, to provide a calibration for the ¹⁴C time scale beyond 30 ka (kilo annum). Improved analytical procedures, and quantitative criteria for sample selection, helped discriminate diagenetically altered samples. The base-line of the calibration curve follows the trend of increasing divergence from calendar ages, as established by previous studies. Superimposed on this trend, four well-defined peaks of excess atmospheric radiocarbon were found ranging in magnitude from 100% to 700%, relative to current levels. They are related to episodes of sea-level rise and reef growth at HP. These peaks appear to be synchronous with Heinrich Events and concentrations of ice-rafted debris found in North Atlantic deepsea cores. Relative timing of sea-level rise and atmospheric ¹⁴C excess imply the following sequence of events: An initial sealevel high is followed by a large increase in atmospheric ¹⁴C as the sea-level subsides. Over about 1800 years, the atmospheric radiocarbon drops to below present ambient levels. This cycle bears a close resemblance to ice-calving episodes of Dansgaard-Oeschger and Bond cycles and the slow-down or complete interruption of the North Atlantic thermohaline circulation. The increases in the atmospheric ¹⁴C levels are attributed to the cessation of the North Atlantic circulation.

INTRODUCTION

Calibration of the radiocarbon time scale over the last ~11 ka has been secured using tree-ring chronology (Becker and Kromer 1993). Data beyond this period, however, are not extensive (Bard 1998). ¹⁴C dating and calendar ages (²³⁰Th/²³⁴U dating) of corals can be determined directly (Bard et al. 1998), an advantage that is particularly relevant for time periods beyond ~25 ka. Coral samples from drill-cores offshore from Barbados (Bard et al. 1990a, 1990b, 1993, 1998) and from Huon Peninsula (Edwards et al. 1993) have been dated with accelerator mass spectrometry (AMS ¹⁴C; Nelson et al. 1977) and thermal ionization mass spectrometry (TIMS, Th/U; Edwards et al. 1987). Where there is an overlap in calendar ages determined from tree-rings and U/Th dating, the two methods are in good agreement (Bard et al. 1998; Kromer and Spurk 1998).

We have extended the previous coral-based ¹⁴C calibration from 30 ka to beyond 50 ka using mass spectrometric ¹⁴C and Th/U dating techniques on corals collected from Reefs II and III at Huon Peninsula, Papua New Guinea (Figure 1). In this paper, we report the new results and discuss possible causes of large excursions of atmospheric ¹⁴C recorded in these corals. Some Th/U ages were previously available for a subset of the same corals (Chappell et al. 1996).

MATERIALS AND METHODS

Most of the corals used in this study were collected from Bobongara (Bobo) and Kanzarua (Kanz) sections that have some of the highest uplift rates at HP (Figure 1). A number of Th/U dates and descriptions of sample locations have been provided by Chappell et al. (1996). General TIMS Useries dating methods were described by Edwards et al. (1987), and procedures used in the present work were previously documented by Stirling et al. (1995) and Esat, (1995). For ¹⁴C dating, CO₂ extracted from corals was converted to graphite, in the presence of H₂, using Fe powder as a catalyst

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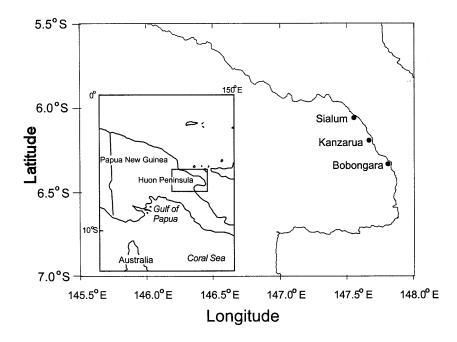


Figure 1 Location of Kanzarua and Bobongara reef sections at Huon Peninsula, Papua New Guinea.

(Vogel et al. 1984), and is described below in detail. Particulars of AMS dating at the Australian National University can be found in Fifield et al. (1993).

The graphitization method employed in the present study reliably produces high graphite yields (>85-95%). Water is formed as the byproduct of the reaction:

$$2H_2 + CO_2 \stackrel{\Rightarrow}{\underset{625 \text{ °C}}{\Rightarrow}} C + 2H_2O$$

$$(1)$$

The addition of dry-ice cold traps attached to the reaction vessel, efficiently trapped water vapor away from the sample and even higher graphite yields (~100%) could be achieved.

To obtain approximately 1–1.5 mg of graphite, ~1.2–1.8 mg of pre-reduced 325 mesh iron powder was typically used resulting in C:Fe ratio of 1 to 1.2. A weighed vycor-glass thimble cup (6 mm diameter and ~1 cm long) containing the iron powder was placed in a larger diameter vycor-glass tube (9 mm diameter and ~15 cm long) and labelled. All vycor glassware was precleaned ultrasonically with distilled water and baked overnight at 900 °C. The reaction tube and connecting pieces were cleaned in 1:1 HNO₃ and rinsed with distilled water, then stored in an oven at ~60 °C until needed. Phosphoric acid (85%) was concentrated on a hotplate with an equal amount of $\rm H_2O_2$ (30%) to oxidize and remove organic impurities.

To reduce the possibility of contamination during the CO₂-graphite conversion we strived to exclude all sources of organic contamination. For this purpose, an all-metal high-vacuum system was constructed with strict control protocol on its use, including the type and age of samples allowed to be processed. The conversion line was made using 0.25-inch stainless steel tubing with Swage-Lock®

joints and stainless-steel bellow valves. All of the component parts were cleaned to remove organic residues and baked under vacuum for extended periods. The line is evacuated using a turbo-molecular pump with a liquid nitrogen trap to prevent oil contamination from the pump. A vacuum of better than 10^{-4} Torr can be maintained during sample processing.

The vacuum was monitored using two thermocouple gauges. Quantitative CO_2 and H_2 pressures were measured by two pressure transducers, Baratron®, which have accuracies of $\pm 0.05\%$ and $\pm 0.5\%$. The more precise Baratron®, was used for low-pressure measurements to monitor the yield of gases evolved from the sample, within a calibrated volume, and the other Baratron®, which has a higher pressure capacity, was used to monitor H_2 pressure during the reduction of Fe powder. The combined volume of the Baratron®, sample-gas storage vessel, and connecting spaces was calibrated using CO_2 produced from precisely weighed amounts of ANU sucrose.

Coral samples were inspected under an optical microscope. For each sample, two thin sections were prepared along the coral longitudinal and latitudinal growth axes. In general, secondary alteration was present at the interstices but not within wall structures.

Calcite content of each sample was determined from X-ray powder diffraction (XRD) spectra. The measurements utilised Cu-Kα radiation (40 kV, 30 mA) and a scintillation counter. The scan rate was 2°2" per minute. Approximately 300 mg of powdered coral was used for each test. The detection limit for calcite was about 2% (Klug and Alexander 1974).

Precleaning of Samples

Corals were inspected under a magnifier and physically cleaned using a dental drill to remove interstitial structures other than walls. The samples were cleaned with distilled water in an ultrasonic bath. They were then dried in an oven (\sim 40–60 °C) and weighed. The minimum required amount of CaCO₃ to yield 1 mg of carbon is approximately 8.5 mg. However, we used larger sample sizes (200–300 mg), and step-wise selective dissolution.

CO₂ Production and Purification

In the present study, samples in the age range 30–60 ka were collected from Reef II and Reef III, including sub-reefs, at Huon Peninsula (Chappell et al. 1996). In addition to textural investigations and XRD measurements, all samples were partially dissolved in several steps, but in a much more severe fashion compared to that previously attempted by Burr et al. (1993). Partial dissolution was first attempted by Bard et al. (1990c); however, its effectiveness was systematically demonstrated later by Burr et al. (1993) using corals older than 70 ka from Barbados and Huon Peninsula. They performed XRD measurements, bulk sample ¹⁴C dating, and ¹⁴C dating following step-wise dissolution. They concluded that the first 50% of acid etching effectively removed secondary, younger age carbon (Burr et al. 1993). In the present study, the first 50% of the coral was etched and discarded, and the rest of the sample was dissolved in three or more aliquots. Each aliquot was AMS dated so that at least three AMS ¹⁴C ages were produced for every sample (Table 1).

Generally, contamination comes from extraneous material containing younger or modern carbon, which shifts "true" ¹⁴C ages to younger values. This could occur not only during the laboratory process, but also in the natural environment through diagenesis and recrystallization. Step-wise selective dissolution could thus help to identify contaminated samples.

The present reaction vessel was designed to facilitate selective dissolution. It consisted of three compartments, the main chamber held the sample, and two accessory branches contained phosphoric acid.

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Table 1 Radiocarbon results

| Lab codea | 0 1 | Dissolved | MC | ¹⁴ C age ^b | Errorc | Errorc | Calcited |
|-----------|------------|-----------|------------------------------------|----------------------------------|------------------|--------|----------|
| (ANUA-) | Sample | percent | pMC | (BP) | (year) | (year) | % |
| 3607 | LIG coral | 0–17 | 0.74 ± 0.13 | 39,460 | +1500 | -1270 | |
| 3608 | (Favid) | 26-30 | 0.25 ± 0.05 | 48,230 | +1880 | -1520 | |
| 3609 | | 34–49 | 0.21 ± 0.03 | 49,450 | +1310 | -1120 | |
| 3610 | | 50-53 | 0.28 ± 0.05 | 47,250 | +1520 | -1280 | |
| 3611 | | 65-70 | 0.21 ± 0.04 | 49,490 | +1660 | -1380 | |
| 3612 | | 71-100 | 0.26 ± 0.04 | 47,970 | +1200 | -1040 | |
| OZD-535 | | 0-100 | 0.52 ± 0.18 | 42,300 | +3410 | -2390 | 2.0 |
| 10017 | Tridacna A | 0-17.2 | 1.03 ± 0.08 | 36,760 | +640 | -600 | |
| 10018 | | 17.2-30.5 | 0.61 ± 0.07 | 41,020 | +980 | -870 | |
| 10019 | | 45.6-73.2 | 0.38 ± 0.05 | 44,830 | +1150 | -1010 | |
| 10020 | | 73.2-100 | 0.27 ± 0.04 | 47,420 | +1270 | -1090 | 0 |
| 9928 | Tridacna B | 50.1-67.2 | 0.31 ± 0.05 | 46,350 | +1340 | -1150 | |
| 9929 | | 67.2-80.5 | 0.34 ± 0.06 | 45,710 | +1590 | -1330 | |
| 9930 | | 80.5-100 | 0.29 ± 0.05 | 46,880 | +1500 | -1260 | 0 |
| 6105 | Kanz 4 | 63.8-76.1 | 1.50 ± 0.13 | 33,730 | +550 | -520 | |
| 6104 | (Favid) | 76.1-88.2 | 1.51 ± 0.12 | 33,660 | +440 | -410 | |
| 6103 | , , | 88.2-100 | 1.55 ± 0.13 | 33,480 | +540 | -500 | |
| OZD-532 | | 0-100 | 2.98 ± 0.21 | 28,200 | +590 | -550 | 0 |
| 4206 | Kanz 9 | 50.8-65.0 | 0.44 ± 0.09 | 43,630 | +1910 | -1540 | |
| 4207 | (Favid) | 65.0-79.0 | 0.41 ± 0.05 | 44,160 | +950 | -850 | |
| 4208 | , | 79.0-100 | 0.49 ± 0.12 | 42,670 | +2320 | -1800 | |
| OZD-533 | | 0-100 | 1.05 ± 0.20 | 36,600 | +1700 | -1400 | 1.0 |
| 7525 | Kanz 11 | 62.3-72.9 | 1.07 ± 0.06 | 36,490 | +490 | -460 | |
| 7526 | (Porites) | 72.9-86.3 | 1.26 ± 0.06 | 35,150 | +420 | -400 | |
| 7527 | , | 86.3-100 | 1.36 ± 0.07 | 34,520 | +450 | -420 | 3.3,1.5 |
| 7528 | Kanz 13 | 47.1-63.0 | 0.84 ± 0.05 | 38,390 | +510 | -480 | , |
| 7529 | (Porites) | 63.0-80.3 | 0.77 ± 0.05 | 38,730 | +560 | -520 | |
| 7530 | , | 80.3-100 | 0.77 ± 0.04 | 39,050 | +450 | -430 | 0.9 |
| 8202 | Kanz 15 | 60.3-73.9 | 0.98 ± 0.09 | 37,120 | +790 | -720 | |
| 8203 | (Porites) | 73.9-85.9 | 0.79 ± 0.11 | 38,890 | +1220 | -1060 | |
| 8204 | , | 85.9-100 | 1.07 ± 0.14 | 36,460 | +1140 | -1000 | 0 |
| 6915 | Kanz U8 | 48.0-64.5 | 1.49 ± 0.11 | 33,800 | +620 | -570 | |
| 6916 | (Porites) | 64.5-81.8 | 1.47 ± 0.11 | 33,930 | +620 | -580 | |
| 6917 | , | 81.8-100 | 1.63 ± 0.12 | 33,090 | +600 | -550 | 1.9 |
| 4230 | Kanz U9 | 46.8-58.6 | 1.63 ± 0.17 | 33,080 | +860 | -780 | |
| 4231 | (Favid) | 58.6-72.1 | 1.24 ± 0.14 | 35,280 | +910 | -840 | |
| 4303 | (, | 72.1-100 | 1.28 ± 0.10 | 35,010 | +750 | -630 | 4.5, 1.8 |
| 4316 | Kanz U10 | 38.6-52.4 | 2.33 ± 0.21 | 30,190 | +680 | -750 | , |
| 4312 | (Porites) | 52.4-83.2 | 1.65 ± 0.19 | 32,980 | +980 | -870 | |
| 4307 | (| 83.2–100 | 1.39 ± 0.11 | 34,350 | +660 | -610 | 2.2 |
| 6908 | Kanz U11 | 46.3–60.0 | 2.30 ± 0.16 | 30,290 | +600 | -560 | |
| 6909 | (Favid) | 60.0–73.7 | 2.38 ± 0.10 2.38 ± 0.17 | 30,040 | +590 | -550 | |
| 6910 | (1 4,14) | 73.7–100 | 2.50 ± 0.17 2.51 ± 0.17 | 29,600 | +580 | -540 | 0 |
| OZD-536 | | 0–100 | 4.33 ± 0.17 | 25,200 | +420 | -400 | U |
| OLD-330 | | 0-100 | 4.33 ± 0.22 | 23,200 | +4 20 | -400 | |

Table 1 Radiocarbon results (Continued)

| Table 1 Radiocarbon results (Continued) | | | | | | | |
|---|-----------|-----------|-----------------|----------------------------------|--------------------|--------------------|----------|
| Lab codea | | Dissolved | | ¹⁴ C age ^b | Error ^c | Error ^c | Calcited |
| (ANUA-) | Sample | percent | pMC | (BP) | (years) | (years) | % |
| 4317 | Kanz U12 | 49.6-61.2 | 2.10 ± 0.19 | 31,050 | +760 | -690 | |
| 4308 | (Favid) | 92.0-100 | 1.84 ± 0.14 | 32,100 | +650 | -600 | 0.5 |
| OZD-537 | | 0-100 | 2.73 ± 0.23 | 28,900 | +710 | -650 | |
| 6829 | Kanz U13 | 50.5-61.4 | 1.84 ± 0.12 | 32,100 | +530 | -500 | |
| 4313 | (Favid) | 61.4–71.9 | 1.07 ± 0.11 | 36,440 | +880 | -790 | |
| 4309 | | 71.9–100 | 1.78 ± 0.17 | 32,350 | +800 | -730 | 0.9 |
| OZD-542 | | 0-100 | 1.91 ± 0.17 | 31,800 | +750 | -680 | |
| 6107 | Kanz U14 | 69.3–79.9 | 1.05 ± 0.10 | 36,630 | +800 | -730 | |
| 6108 | (Porites) | 79.9–89.1 | 0.83 ± 0.07 | 38,500 | +700 | -640 | |
| 6109 | | 89.1–100 | 1.04 ± 0.08 | 36,600 | +680 | -630 | 2.5 |
| OZD-540 | | 0–100 | 1.53 ± 0.21 | 33,600 | +1200 | -1030 | |
| 6825 | Kanz U15 | 70.0–80.5 | 0.97 ± 0.07 | 37,210 | +590 | -550 | |
| 6824 | (Favid) | 80.5–89.5 | 0.94 ± 0.07 | 37,490 | +600 | -560 | |
| 6823 | | 89.5–100 | 1.22 ± 0.07 | 35,410 | +500 | -470 | 0.1 |
| 6828 | Kanz U16 | 59.0-73.1 | 1.91 ± 0.10 | 31,810 | +450 | -420 | |
| 6827 | (Porites) | 73.1–86.7 | 2.04 ± 0.11 | 31,290 | +430 | -410 | |
| 6826 | | 86.7–100 | 1.76 ± 0.11 | 32,450 | +510 | -480 | 3.2 |
| 6818 | Kanz A | 47.9–58.2 | 2.96 ± 0.14 | 28,280 | +380 | -360 | |
| 6817 | (Favid) | 58.2-69.0 | 2.81 ± 0.16 | 28,690 | +480 | -450 | |
| 6816 | | 69.0–100 | 2.90 ± 0.16 | 28,440 | +470 | -440 | 1.0 |
| 4209 | Bobo U10 | 56.2-67.7 | 2.47 ± 0.16 | 28,730 | +550 | -510 | |
| 4210 | (Favid) | 67.7–97.4 | 1.62 ± 0.12 | 33,120 | +640 | -590 | 0 |
| 4212 | Bobo U11 | 51.9–61.7 | 3.46 ± 0.22 | 27,020 | +530 | -500 | |
| 4213 | (Favid) | 61.7–71.5 | 4.34 ± 0.34 | 25,210 | +660 | -610 | |
| 4214 | | 71.5–100 | 4.66 ± 0.36 | 24,630 | +640 | -590 | 7.5, 2.9 |
| 4315 | Bobo U17 | 44.5-53.1 | 3.05 ± 0.26 | 28030 | +710 | -650 | |
| 4216 | (Favid) | 53.1-63.9 | 4.64 ± 0.35 | 24,670 | +640 | -590 | |
| 4217 | | 63.9–100 | 3.65 ± 0.24 | 26,590 | +550 | -520 | 5.1, 2.4 |
| OZD-538 | | 0-100 | 5.24 ± 0.30 | 23,700 | +470 | -450 | |
| 9507 | Bobo U18 | 0-31.0 | 6.09 ± 0.25 | 22,480 | +340 | -320 | |
| 4218 | (Favid) | 45.9–60.0 | 1.57 ± 0.16 | 33,340 | +880 | -800 | |
| 4219 | | 60.0-83.5 | 1.54 ± 0.21 | 33,530 | +1160 | -1010 | |
| 4220 | | 83.5-100 | 1.89 ± 0.14 | 31,880 | +600 | -560 | 0 |
| OZD-544 | | 0–100 | 1.84 ± 0.20 | 22,100 | +930 | -830 | |
| 4221 | Bobo U20 | 49.8–63.3 | 2.27 ± 0.22 | 30,410 | +810 | -740 | |
| 4222 | (Favid) | 63.3–78.1 | 2.13 ± 0.19 | 30,920 | +750 | -680 | |
| 4223 | | 78.1–100 | 2.06 ± 0.17 | 31,190 | +690 | -630 | 0 |
| 4225 | Bobo U21 | 56.8-73.1 | 3.65 ± 0.30 | 26,590 | +680 | -630 | |
| 4226 | (Favid) | 73.1–100 | 3.91 ± 0.30 | 26,050 | +640 | -590 | 1.0 |
| OZD-543 | | 0-100 | 7.08 ± 0.32 | 21,250 | +370 | -340 | |
| 4227 | Bobo U24 | 49.5-61.3 | 3.30 ± 0.27 | 27,400 | +680 | -630 | |
| 4228 | (Favid) | 61.3-72.9 | 2.92 ± 0.25 | 28,390 | +710 | -650 | |
| 4229 | | 72.9-100 | 3.44 ± 0.25 | 27,070 | +600 | -560 | 1.0 |
| 9504 | Bobo U28 | 0-30.0 | 18.2 ± 0.62 | 13,700 | +280 | -270 | |

Table 1 Radiocarbon results (Continued)

| Lab code ^a | | Dissolved | | ¹⁴ C age ^b | Errorc | Errorc | Calcited |
|-----------------------|-----------|-----------|------------------|----------------------------------|---------|---------|----------|
| (ANUA-) | Sample | percent | pMC | (BP) | (years) | (years) | % |
| 6912 | (Favid) | 60.3-71.8 | 4.71 ± 0.31 | 24,550 | +540 | -510 | |
| 6913 | | 71.8-86.7 | 4.53 ± 0.30 | 24,870 | +550 | -510 | |
| 6914 | | 86.7-100 | 5.34 ± 0.35 | 23,540 | +550 | -520 | 2.0 |
| OZD-539 | | 0-100 | 7.91 ± 0.62 | 20,400 | +330 | -320 | |
| 9502 | Bobo U30 | 0-29.0 | 7.85 ± 0.43 | 20,440 | +450 | -430 | |
| 6115 | (Favid) | 64.8-76.0 | 4.09 ± 0.26 | 25,680 | +530 | -500 | |
| 6114 | | 76.0-87.3 | 4.10 ± 0.27 | 25,660 | +540 | -510 | |
| 6113 | | 87.3-100 | 3.70 ± 0.24 | 26,500 | +550 | -510 | 1.0 |
| OZD-541 | | 0-100 | 7.75 ± 0.27 | 20,550 | +280 | -280 | |
| 8205 | SEN(N)8 | 62.3-72.0 | 0.48 ± 0.09 | 42,970 | +1680 | -1390 | |
| 8206 | (Porites) | 72.0-86.9 | 0.60 ± 0.08 | 41,130 | +1160 | -1010 | |
| 8203 | | 86.9-100 | 0.59 ± 0.10 | 41,220 | +1500 | -1260 | 1.6 |
| 6822 | GBR-A | 62.5-74.4 | 17.06 ± 0.75 | 14,200 | +360 | -340 | |
| 6820 | (Galaxea | 74.4-86.7 | 16.93 ± 0.72 | 14,270 | +350 | -340 | |
| 6819 | clavus) | 86.7–100 | 16.67 ± 0.64 | 14,390 | +320 | -300 | 0 |

^aAll samples were measured at ANU except the OZD-series which were measured at ANSTO (Australian Nuclear Science and Technology Organisation).

The vessel was connected by a Cajon Ultra-Torr® fitting to the main vacuum line. After 1.5 hr of pumping and out-gassing of the acid, a vacuum of 10^{-4} Torr could be achieved. The vessel was then tilted to drip acid onto the sample to start the reaction. The CO_2 produced was dried by passing through two spiral traps immersed in a dry ice-ethanol mixture (-78.8 °C), and cryogenically collected at liquid nitrogen temperature. The CO_2 pressure was measured at intervals to determine the amount of sample dissolved. The dried CO_2 was then transferred via one of several outlets to the graphitisation apparatus described previously. The iron powder had previously been reduced in situ with ~ 400 Torr H_2 at 400 °C for several hours. The setup includes a cold finger to trap water vapor, and a pressure transducer to monitor the progress of the conversion reaction. The apparatus has a double valve system for isolating and removing the samples from the vacuum system. ANU sucrose was used as a standard. The sucrose was weighed and placed in a vycor tube (6 mm diameter and ~ 20 cm long) with CuO and Ag wire. The tube was evacuated and sealed with a torch and placed in an oven at 900 °C for 8–10 hr to combust and convert the sucrose to CO_2 . The rest of the procedure was the same as that used for samples.

After evacuating the sample tube ($<10^{-4}$ Torr) at liquid nitrogen temperature, it was filled with approximately 2.2 times molar fraction of excess H_2 relative to CO_2 . The valves were closed and graphitization reaction started by inserting the end of the tube into a high temperature oven at 625 °C at a separate location. Up to four sample tubes could be heated at the same time. Water vapor produced by the reaction was removed by a cold finger in a dry-ice-ethanol bath attached to the graphitization apparatus. The change in pressure due to H_2 consumption was monitored with pressure transducer. Pressure readings indicated that most of the reaction had occurred during he first 0.5–1 hr, and

^bAge obtained using Libby's half life (5568 yr).

^cErrors are quoted as two sigma.

^dXRD analyses for bulk samples. Where two entries are given, the second corresponds to wall fraction following mechanical cleaning to remove septa.

was essentially complete within 3 hr. However, to ensure full conversion, the process was continued for at least 6 hr.

Graphite yields were calculated by both weighing the residue, and from the measured pressure difference, and ranged from 85% to 100%. Conversion efficiency was typically greater than 95%. At such high efficiencies, graphitization related carbon isotope fractionation is expected to be minimal (e.g. McNichol et al. 1992; Kitagawa et al. 1993; Brown and Southon 1997).

Background ¹⁴C levels were monitored regularly using the international old calcite standard, IAEA-C1 (Rozanski et al. 1992). The background values ranged from 0.04 pMC (percent modern carbon) to 0.30 pMC, and were typically 0.25 pMC, equivalent to an age of 48,000 yr. Therefore, samples as old as 40,000 yr could be reliably measured using the present system and procedures.

RESULTS AND DISCUSSION

Integrity of Samples

The basic requirement in ¹⁴C and U/Th dating is the assumption of closed system behavior for the isotopes used in the measurement. In general, two sources of contamination can be expected in ¹⁴C dating. One from contamination in the laboratory, including during the graphitization process, and the other is from inherent diagenesis in corals.

The former can be assessed by using "old" carbonate, which ideally contains no 14 C. In our laboratory, we have used Carrara marble (IAEA C-1) that yielded a range of ages from 47,000 to 50,000 BP for the total analytical blank, including AMS analysis. Therefore, any dates younger than this age are reliable, assuming sample integrity. Numerous studies have attempted to characterize the cause and circumstances of diagenetic processes in corals (e.g. Bathurst 1968; Lighty 1985; Moore 1989; Strasser et al. 1997) however, the nature of the mechanism is still unclear (Bar-Matthews et al. 1993). To evaluate its effects in corals several criteria have previously been employed such as microscopic textural analysis, XRD measurements, stable isotope measurements, total uranium contents, the amount of 232 Th in the coral, trace element analysis, and measurement of the initial 234 U/ 238 U ratio δ^{234} U(T).

Textural Investigation and XRD Analysis

Evaluating the possibility of open system behavior for ¹⁴C is much more difficult than for Th and U isotopes. However, to some extent, it is possible to check for secondary material, as pristine coral consists only of aragonite. XRD can be used to check for secondary calcite. We have investigated textures with a petrographic microscope and XRD analyses. In most of the samples, secondary fillings were observed in the voids of the coral. These were mechanically removed using a dental drill prior to leaching. Thus, even in cases where minor secondary replacement was observed, and detected in XRD analyses of whole samples, they were removed. All the samples had typically less than 2% calcite, which is within the uncertainty of detection limits (Table 1).

Stable Isotope Analysis

For each sample, stable isotopes of oxygen and carbon were measured. If marine carbonates preserve their original composition, both isotopes values should be around zero. Stein et al. (1993) conducted an extensive study of Huon LIG corals. Oxygen isotopes are likely to have been affected if exotic isotopes were introduced into the coral skeleton through percolation of meteoric water. However, oxygen and carbon isotopes appear not to have responded to small-scale diagenetic alteration

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in corals. The present corals are significantly younger than LIG corals used by Stein et al. (1993), and as such they are less likely to have detectable variations in oxygen or carbon isotopes correlated with small scale diagenesis. The test results showed that carbon and oxygen isotope values were within acceptable ranges as found by Stein et al. (1993).

Uranium Concentration

Uranium concentration of corals in this study range between ~2 and 3.5 ppm. These values are consistent with previously reported results for younger age corals (Hamelin et al. 1991), Holocene corals (Edwards et al. 1988; Zhu et al. 1993), and the LIG corals (Edwards et al. 1986; Hamelin et al. 1991; Edwards et al. 1987; Stirling et al. 1995, 1998). Species dependence of uranium content in corals has been reported previously (Zhu et al. 1993; Stirling 1995). Stirling (1995) reported that uranium content of *Acropora* was higher than other species such as *Porites* or *Favidae*, but Zhu et al. (1993), pointed out that *Acropora* and *Montipora* corals have inherently higher ²³⁸U values in Holocene corals. Therefore, the variations in Uranium content might be species dependent, but generally range from 2 to 4 ppm. All corals used in the present study are well within this range.

²³²Th Concentration in the Corals

²³²Th in modern corals is typically less than 0.5 ppb (Edwards et al. 1986; Chen et al. 1991). In the freshwater or marine environment, voids can be filled with externally introduced detritus material. Therefore, if ²³²Th abundance departs significantly from the 0.5 ppb limit, it may indicate the presence of secondary contamination. Previous researchers have attempted to correct for excess levels of ²³²Th, although the correction is not generally considered to be reliable (e.g. Chen, et al. 1991). The effect is to shift the measured Th-U age toward younger ages.

In the present study, for most corals, 232 Th is well below the 0.5 ppb limit. One coral has ≈ 1.1 ppb. This coral was collected at a depth of -175 m in the Great Barrier Reef (Veeh and Veevers 1970). It had not been exposed to aerial or subaerial environment. Therefore, the excess 232 Th must have been added in the marine environment. There are no discernable systematic correlations between 232 Th and U/Th age, δ^{234} U(T), or Δ^{14} C.

$\delta^{234}U(T)$

Behavior of trace chemical elements with degree of diagenesis was studied in Caribbean corals by Bar-Matthews et al. (1993). These authors investigated both textural structures with an optical microscope and a scanning electron microscope, and then conducted chemical analyses. They concluded that textural analysis is necessary, yet it is not a quantitative or reliable test in evaluating the degree of diagenesis. Most of the alteration was observed in the voids of corals as secondary filling and calcium carbonate deposition. However, cementation in these corals was not always by calcite, but in many cases through secondary aragonite (Bar-Matthews et al. 1993). Therefore, they were presumably altered in the marine environment, quite likely at the intertidal or supertidal zone as indicated by other studies (e.g. Bar-Mathews et al. 1993; Lighty 1985).

Many studies have detailed the shift in $\delta^{234}U(T)$ through diagenesis (Edwards 1988; Ku et al. 1990; Banner et al. 1991; Bard et al. 1991; Chen et al. 1991; Stein et al. 1993; Gallup et al. 1994; Stirling et al. 1995, 1998). The magnitude of $\delta^{234}U(T)$ tends to be elevated as the degree of diagenesis increases. Gallup et al. (1994) demonstrated this trend for corals from Barbados and the Bahamas. Also, Stirling (1998) observed a similar increase in $\delta^{234}U(T)$ with age for corals from Western Australia. We have used a similar acceptance criterion: $\delta^{234}U(T) = 149 \pm 10\%$ as used in previous studies (e.g. Stirling et al. 1998).

Radiocarbon and Th-U Dating of Corals from Huon Peninsula

Veracity of the results from ¹⁴C and Th-U dating of corals ultimately depend on the quality of the samples. This is a vexing issue that cannot be resolved with complete certainty, but relies on a battery of tests, each of which provide, at best, circumstantial evidence for or against. We have employed a quantitative test of ¹⁴C dates that relies on progressive dissolution of coral samples. We have determined empirically that the first 50% of a leached coral sample is more likely to have been contaminated by extraneous, younger age carbon, than the last 50% (Burr et al. 1993). We have adopted a procedure whereby the first 50% of a sample was selectively dissolved and discarded. The remainder was dissolved in at least three steps and each aliquot was 14C-AMS dated. This procedure was applied to each dated sample. The data for all of the samples can be found in Figure 2 and Table 1. In every case, within twice the statistical uncertainty, the last three aliquots have identical ¹⁴C ages. Samples included in the discussion below pass all of these tests within limits of assigned errors. A major concern with ¹⁴C dates for samples older than 30 ka is the possibility of contamination with younger age or modern carbon. There is no absolute a priori test to determine the issue. In particular, for the 52 ka samples reported below, the ¹⁴C ages range from 35 to 45 ka BP and may need to be viewed with some caution. However, we consider that the selective dissolution test is the best quantitative indicator that can be devised at present.

We tested the effectiveness of this procedure by using corals as old as 120,000 yr that grew during the Last Interglacial (LIG). Burr et al. (1992), conducted step-wise dissolution of a reef IV coral, from Huon Peninsula, which was previously dated (72.8 \pm 2.2 ka; Chappell et al. 1996). The first 40% of this coral was contaminated by younger carbon; the next 60% gave an age close to the analytical background. We have used a LIG coral from Western Australia (WA-1) that was previously Th-U dated (Stirling et al. 1995) and has an age of 124.7 \pm 0.9 ka with an initial δ^{234} U(T) ratio of 152 \pm 2. WA-1 appears to be pristine under the optical microscope, and no calcite was detected by XRD analysis. The first 0–17% aliquot from WA-1 shows a considerably younger ¹⁴C age compared to the analytical background (~48,000 BP). However, the next 5 aliquots, indicate an almost uniform age of around 49,000 BP. Two other LI corals from Western Australia were treated similarly and radiocarbon dated. For the first coral, the (0% to 54%) aliquot gave an age of 35,270 \pm 1490 BP. The remainder of the sample (54% to 100%) was dated at 47,690 \pm 1500 BP. The second coral was dissolved in three steps (15% to 28%, 28% to 36%, and 55% to 100%) and yielded the corresponding ages of 40,960 BP, 42,560 BP, and 50,770 BP, respectively.

Giant clams, *Tridacna gigas* do not incorporate U in their aragonite skeleton and cannot be dated by U-series methods. *Tridacna* were previously collected from Huon terraces IIIa and IIIb from locations of known age. The harder skeleton of *Tridacna* may be expected to be resistant to alteration compared to fossil corals and therefore provide more reliable ¹⁴C ages. However, the estimated age of available *Tridacna* samples in the collection was over 50,000 yr. We ¹⁴C dated two samples. A *Tridacna* from reef IIIa shows a nearly constant age profile at ~47,000 BP which is close to the analytical background of the present system. Another *Tridacna* was treated in a similar fashion. The first 17% indicated younger carbon contamination. However the last ~50% yielded consistent results of ~47,000 BP. Both *Tridacna* consist of 100% aragonite as determined by XRD analysis. It appears that the harder skeleton of *Tridacna* is also susceptible to alteration in the outer layers.

Most of the samples used in the present study have also been measured at another laboratory. Samples were sent to ANSTO (Australian Nuclear Science and Technology Organization) for ¹⁴C dating (Hotchkis et al. 1996). Measurements of ¹⁴C at ANSTO were on whole coral samples, without selective dissolution. This is shown as the (0–100%) fully dissolved fraction in Figure 2. As can be seen,

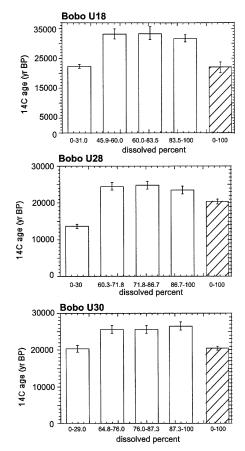


Figure 2a All the samples used in the present study were subjected to step-wise dissolution as shown in these examples. The fraction of dissolved coral was determined from pressure and temperature measurements of evolved CO₂ in a calibrated volume. The hatched columns (0–100%) show the results of bulk analysis for a different piece of the same coral.

all of the dates measured at ANSTO are significantly younger than those for partially dissolved samples. This, and the discussion in previous sections, clearly indicate the importance of step-wise dissolution.

In summary, the above experiments illustrate the following points: First, the necessity of strict criteria based on textural analysis, XRD and step-wise dissolution. Second, that approximately the first 20% of dissolved samples are likely to be contaminated by younger carbon. And third, that after 50% dissolution, most of the contamination is eliminated.

Overall, tests using step-wise dissolution appear to be the only way to ascertain the veracity of ¹⁴C ages of corals. Our results are consistent with the experiments of Burr et al. (1993) that after 50% etching, ¹⁴C age approaches analytical background levels. The criteria we have applied to each sample requires ¹⁴C dating of at least three aliquots, after the first 50% of the sample is selectively dissolved and discarded. The results were considered acceptable if the three dates agreed within analytical uncertainties. Although, this appears to be the best quantitative method available so far, it does not guarantee that samples are fully free of diagenesis at a level not to affect age determinations.

The new data comprise of 25 new TIMS Th-U dates and over 200 AMS available dates (Table 2). The ¹⁴C ages and Th-U calendar ages are shown in Figure 3 together with data from other studies. Sample age was calculated from the weighted mean of three available ages measured following 50%

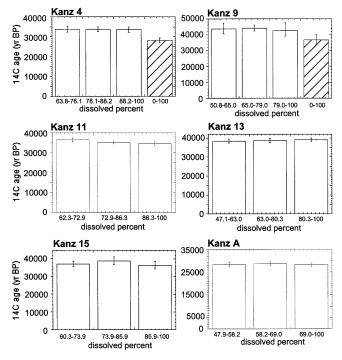


Figure 2b Same as Figure 2a but for different corals

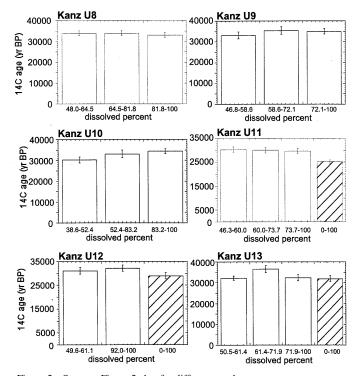


Figure 2c $\,$ Same as Figure 2a but for different corals

Table 2 U/Th and ¹⁴C results

| $\begin{array}{c ccccccccccccccccccccccccccccccccccc$ | Table 2 U/TII | and Cresuits | | | | |
|---|-----------------------|----------------|----------------------|-----------------|-------|-----------------|
| Kanzarua Kanz 4 51.5 ± 1.6 33.6 ± 0.4 6678 IIIa -58 (86) Kanz 4 51.2 ± 0.8 1759 IIIa -72 (78) Kanz 9 53.4 ± 0.3 43.9 ± 0.8 1759 IIIa -72 (78) Kanz 19 54.6 ± 0.7 54.6 ± 0.7 54.6 ± 0.7 IIIb -59 (48) Kanz 13 38.3 ± 0.5 38.9 ± 0.3 -187 IIIb -59 (48) Kanz 15 42.0 ± 0.6 37.5 ± 0.6 511 IIIc -79 (39) Kanz U8 37.0 ± 0.6 33.6 ± 0.4 333 IIa -77 (27) Kanz U9 42.3 ± 0.2 34.7 ± 0.5 1075 IIa -91 (28) Kanz U9 42.3 ± 0.2 34.7 ± 0.5 1075 IIa -91 (28) Kanz U9 42.3 ± 0.3 33.2 ± 0.5 2162 IIIc -73 (49) Kanz U10 9 43.7 ± 0.3 33.2 ± 0.5 2162 IIIc -73 (49) Kanz U11 37.9 ± 0.3 30.0 ± 0.3 1351 IIa -79 (27) Kanz U12 37.5 ± 0.7 31.7 ± 0.5 794 IIa -74 (26) | | U/Th age | ¹⁴ C age | Δ^{14} C | | Sea-levels |
| Kanz 4 51.5 ± 1.6 33.6 ± 0.4 6678 IIIa $-58 (86)$ Kanz 9 53.4 ± 0.3 43.9 ± 0.8 1759 IIIa $-72 (78)$ Kanz 9T 54.6 ± 0.7 Kanz 11P 51.3 ± 0.3 35.4 ± 0.3 5063 IIIa $-65 (78)$ Kanz 13P 38.3 ± 0.5 38.9 ± 0.3 -187 IIIb $-59 (48)$ Kanz 15P 42.0 ± 0.6 37.5 ± 0.6 511 IIIc $-79 (39)$ Kanz U8P 37.0 ± 0.6 33.6 ± 0.4 333 IIa $-77 (27)$ Kanz U9 42.3 ± 0.2 34.7 ± 0.5 1075 IIa $-91 (28)$ Kanz U9T 42.2 ± 0.3 33.2 ± 0.5 2162 IIIc $-73 (49)$ Kanz U10P 43.7 ± 0.3 33.2 ± 0.5 2162 IIIc $-73 (49)$ Kanz U11 37.9 ± 0.3 30.0 ± 0.3 1351 IIa $-79 (27)$ Kanz U14T 34.8 ± 0.3 | Samplea | (ka cal BP) | (ka BP) ^b | (‰)c | Reefd | and height (m)e |
| Kanz 4^{α} 51.2 ± 0.8 Kanz 9 53.4 ± 0.3 43.9 ± 0.8 1759 IIIa -72 (78) Kanz 9 ^T 54.6 ± 0.7 Kanz 11P 51.3 ± 0.3 35.4 ± 0.3 5063 IIIa -65 (78) Kanz 13P 38.3 ± 0.5 38.9 ± 0.3 -187 IIIb -59 (48) Kanz 15P 42.0 ± 0.6 37.5 ± 0.6 511 IIIc -79 (39) Kanz U9P 42.3 ± 0.2 34.7 ± 0.5 1075 IIa -91 (28) Kanz U9Φ 41.8 ± 0.6 Kanz U9Φ 42.3 ± 0.2 34.7 ± 0.5 1075 IIa -91 (28) Kanz U9T 42.2 ± 0.3 Kanz U10P 43.7 ± 0.3 33.2 ± 0.5 2162 IIIc -73 (49) Kanz U11 37.9 ± 0.3 30.0 ± 0.3 1351 IIa -79 (27) Kanz U12 37.5 ± 0.7 31.7 ± 0.5 794 IIa -79 (26) Kanz U13 35.8 ± 0.4 33.8 ± 0.5 139 IIa -74 (26) Kanz U14 34.8 ± 0.8 37.4 ± 0.4 -360 IIa -72 (26) Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II -71 (23) Kanz U16P 30.4 ± 0.4 31.8 ± 0.3 -249 II -56 (29) Kanz U16P 30.4 ± 0.4 31.8 ± 0.3 -249 II -56 (29) Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II -80 (25) Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa -74 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc -68 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 33.4 ± 0.6 Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 33.3 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 33.0 ± 0.5 Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -77 (27) Others SEN(N)8P 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA | Kanzarua | | | | | |
| Kanz 9 53.4 ± 0.3 43.9 ± 0.8 1759 IIIa $-72 (78)$ Kanz 9 ^T 54.6 ± 0.7 Kanz 11P 51.3 ± 0.3 35.4 ± 0.3 5063 IIIIa $-65 (78)$ Kanz 13P 38.3 ± 0.5 38.9 ± 0.3 -187 IIIIb $-59 (48)$ Kanz 15P 42.0 ± 0.6 37.5 ± 0.6 511 IIIc $-79 (39)$ Kanz U8P 37.0 ± 0.6 33.6 ± 0.4 333 IIa $-77 (27)$ Kanz U9 42.3 ± 0.2 34.7 ± 0.5 1075 IIa $-71 (27)$ Kanz U9 ^α 41.8 ± 0.6 Kanz U9 ^α 42.2 ± 0.3 Kanz U10P 43.7 ± 0.3 33.2 ± 0.5 2162 IIIc $-73 (49)$ Kanz U11 37.9 ± 0.3 30.0 ± 0.3 1351 IIa $-79 (27)$ Kanz U12 37.5 ± 0.7 31.7 ± 0.5 794 IIa $-79 (26)$ Kanz U13 35.8 ± 0.4 33.8 ± 0.5 139 IIa $-74 (26)$ Kanz U14 34.8 ± 0.8 37.4 ± 0.4 -360 IIa $-72 (26)$ Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II $-71 (23)$ Kanz U16P 30.4 ± 0.4 31.8 ± 0.3 -249 II $-56 (29)$ Kanz U16P 30.4 ± 0.4 31.8 ± 0.3 -249 II $-56 (29)$ Kanz U10 37.5 ± 0.7 28.5 ± 0.3 1696 II $-80 (25)$ Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa $-74 (49)$ Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa $-76 (49)$ Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIc $-86 (20)$ Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc $-68 (30)$ Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc $-68 (30)$ Bobo U21 33.4 ± 0.6 Bob U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc $-76 (30)$ Bobo U24 32.2 ± 0.2 26.0 ± 0.3 791 IIc $-76 (30)$ Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc $-76 (30)$ Bobo U24 32.2 ± 0.2 26.0 ± 0.3 791 IIc $-76 (30)$ Bobo U24 32.2 ± 0.2 26.0 ± 0.3 791 IIc $-77 (27)$ Others SEN(N)8P 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc $-77 (27)$ Others SEN(N)8P 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc $-77 (27)$ Others SEN(N)8P 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc $-77 (27)$ Others | Kanz 4 | 51.5 ± 1.6 | 33.6 ± 0.4 | 6678 | IIIa | -58 (86) |
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| Kanz 11P 51.3 \pm 0.3 35.4 \pm 0.3 5063 IIIa -65 (78) Kanz 13P 38.3 \pm 0.5 38.9 \pm 0.3 -187 IIIb -59 (48) Kanz 15P 42.0 \pm 0.6 37.5 \pm 0.6 511 IIIc -79 (39) Kanz U8P 37.0 \pm 0.6 33.6 \pm 0.4 333 IIa -77 (27) Kanz U9 42.3 \pm 0.2 34.7 \pm 0.5 1075 IIa -91 (28) Kanz U9 $^{\alpha}$ 41.8 \pm 0.6 Kanz U9 $^{\alpha}$ 42.2 \pm 0.3 43.0 \pm 0.3 1351 IIa -79 (27) Kanz U11 37.9 \pm 0.3 30.0 \pm 0.3 1351 IIa -79 (27) Kanz U12 37.5 \pm 0.7 31.7 \pm 0.5 794 IIa -79 (26) Kanz U13 35.8 \pm 0.4 33.8 \pm 0.5 139 IIa -74 (26) Kanz U15 33.4 \pm 0.2 36.7 \pm 0.3 30.4 \pm 0.3 IIa -74 (26) Kanz U15 33.4 \pm 0.2 36.7 \pm 0.3 1696 II -71 (23) Kanz U16 $^{\alpha}$ 30.4 \pm 0.4 31.8 \pm 0.3 -249 II -756 (29) Kanz A 37.5 \pm 0.2 28.5 \pm 0.3 1696 II -80 (25) Bobongara Bobo U10 37.2 \pm 0.2 28.5 \pm 0.3 1696 II -80 (25) Bobo U17 32.0 \pm 0.2 26.6 \pm 0.4 741 IIc -86 (20) Bob U18 42.4 \pm 0.2 32.7 \pm 0.5 1897 IIb -103 (37) Bobo U20 37.6 \pm 0.2 26.3 \pm 0.5 1897 IIb -103 (37) Bobo U21 29.7 \pm 0.2 26.3 \pm 0.5 373 IIc -68 (30) Bob U24 32.2 \pm 0.2 27.6 \pm 0.4 583 IIc -76 (49) Bobo U24 33.4 \pm 0.6 Bob U24 32.2 \pm 0.2 27.6 \pm 0.4 583 IIc -76 (30) Bob U24 32.2 \pm 0.2 27.6 \pm 0.4 583 IIc -76 (30) Bob U24 32.2 \pm 0.2 27.6 \pm 0.4 583 IIc -76 (30) Bob U24 32.2 \pm 0.2 27.6 \pm 0.4 583 IIc -76 (30) Bob U24 32.2 \pm 0.2 27.6 \pm 0.4 583 IIc -77 (27) Others SEN(N)8P 52.4 \pm 1.3 41.7 \pm 0.8 2172 IIIc nA | | 53.4 ± 0.3 | 43.9 ± 0.8 | 1759 | IIIa | -72 (78) |
| Kanz 13 ^p 38.3 ± 0.5 38.9 ± 0.3 -187 IIIb -59 (48) Kanz 15 ^p 42.0 ± 0.6 37.5 ± 0.6 511 IIIc -79 (39) Kanz U8 ^p 37.0 ± 0.6 33.6 ± 0.4 333 IIa -77 (27) Kanz U9 42.3 ± 0.2 34.7 ± 0.5 1075 IIa -91 (28) Kanz U9 ^α 41.8 ± 0.6 Kanz U9 ^π 42.2 ± 0.3 Kanz U10 ^p 43.7 ± 0.3 33.2 ± 0.5 2162 IIIc -73 (49) Kanz U11 37.9 ± 0.3 30.0 ± 0.3 1351 IIa -79 (27) Kanz U12 37.5 ± 0.7 31.7 ± 0.5 794 IIa -79 (26) Kanz U13 35.8 ± 0.4 33.8 ± 0.5 139 IIa -74 (26) Kanz U14 ^T 34.8 ± 0.8 37.4 ± 0.4 -360 IIa -72 (26) Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II -71 (23) Kanz U16 ^p 30.4 ± 0.4 31.8 ± 0.3 -249 II -56 (29) Kanz U16 37.2 ± 0.2 28.5 ± 0.3 1696 II -80 (25) Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa -76 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc -68 (30) Bobo U24 33.4 ± 0.6 Bobo U25 33.4 ± 0.6 Bobo U26 33.4 ± 0.6 Bobo U27 33.4 ± 0.6 Bobo U28 33.4 ± 0.6 Bobo U29 37.6 ± 0.2 26.0 ± 0.3 1388 IIc -76 (30) Bobo U20 37.6 ± 0.2 26.0 ± 0.3 791 IIc -77 (27) Others SEN(N)8 ^p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA | | 54.6 ± 0.7 | | | | |
| Kanz 15 ^p 42.0 ± 0.6 37.5 ± 0.6 511 IIIc $-79 (39)$ Kanz U8 ^p 37.0 ± 0.6 33.6 ± 0.4 333 IIa $-77 (27)$ Kanz U9 42.3 ± 0.2 34.7 ± 0.5 1075 IIa $-91 (28)$ Kanz U9 ^α 41.8 ± 0.6 Kanz U9 ^T 42.2 ± 0.3 Kanz U10 ^p 43.7 ± 0.3 33.2 ± 0.5 2162 IIIc $-73 (49)$ Kanz U11 37.9 ± 0.3 30.0 ± 0.3 1351 IIa $-79 (27)$ Kanz U12 37.5 ± 0.7 31.7 ± 0.5 794 IIa $-79 (26)$ Kanz U13 35.8 ± 0.4 33.8 ± 0.5 139 IIa $-74 (26)$ Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II $-71 (23)$ Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II $-71 (23)$ Kanz U16 ^p 30.4 ± 0.4 31.8 ± 0.3 -249 II $-56 (29)$ Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II $-80 (25)$ Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa $-74 (49)$ Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa $-76 (49)$ Bobo U17 32.0 ± 0.2 32.7 ± 0.5 1897 IIb $-103 (37)$ Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb $-84 (40)$ Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc $-68 (30)$ Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc $-76 (30)$ Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc $-76 (30)$ Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc $-76 (30)$ Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc $-76 (30)$ Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc $-76 (30)$ Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc $-77 (27)$ Others SEN(N)8 ^p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA | Kanz 11 ^p | 51.3 ± 0.3 | 35.4 ± 0.3 | 5063 | IIIa | -65 (78) |
| Kanz U8 ^p 37.0 ± 0.6 33.6 ± 0.4 333 IIa $-77 (27)$ Kanz U9 42.3 ± 0.2 34.7 ± 0.5 1075 IIa $-91 (28)$ Kanz U9 ^α 41.8 ± 0.6 Kanz U9 ^T 42.2 ± 0.3 Kanz U10 ^p 43.7 ± 0.3 33.2 ± 0.5 2162 IIIc $-73 (49)$ Kanz U11 37.9 ± 0.3 30.0 ± 0.3 1351 IIa $-79 (27)$ Kanz U12 37.5 ± 0.7 31.7 ± 0.5 794 IIa $-79 (26)$ Kanz U13 35.8 ± 0.4 33.8 ± 0.5 139 IIa $-74 (26)$ Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II $-71 (23)$ Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II $-71 (23)$ Kanz U16 ^p 30.4 ± 0.4 31.8 ± 0.3 -249 II $-56 (29)$ Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II $-80 (25)$ Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa $-74 (49)$ Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa $-76 (49)$ Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIb $-103 (37)$ Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb $-84 (40)$ Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc $-68 (30)$ Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc $-68 (30)$ Bobo U24 33.4 ± 0.6 Bobo U24 33.4 ± 0.6 Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc $-76 (30)$ Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc $-76 (30)$ Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc $-76 (30)$ Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc $-76 (30)$ Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc $-76 (30)$ Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc $-77 (27)$ Others SEN(N)8 ^p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc IIIc $-77 (27)$ Others SEN(N)8 ^p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc IIIc | | 38.3 ± 0.5 | 38.9 ± 0.3 | -187 | | -59 (48) |
| Kanz U9 42.3 ± 0.2 34.7 ± 0.5 1075 IIa -91 (28) Kanz U9 ^α 41.8 ± 0.6 Kanz U9 ^T 42.2 ± 0.3 Kanz U10 ^p 43.7 ± 0.3 33.2 ± 0.5 2162 IIIc -73 (49) Kanz U11 37.9 ± 0.3 30.0 ± 0.3 1351 IIa -79 (27) Kanz U12 37.5 ± 0.7 31.7 ± 0.5 794 IIa -79 (26) Kanz U13 35.8 ± 0.4 33.8 ± 0.5 139 IIa -74 (26) Kanz U14 ^T 34.8 ± 0.8 37.4 ± 0.4 -360 IIa -72 (26) Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II -71 (23) Kanz U16 ^p 30.4 ± 0.4 31.8 ± 0.3 -249 II -56 (29) Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II -80 (25) Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa -74 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc -68 (30) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc -68 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -77 (27) Others SEN(N)8 ^p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA GBR-A 17.0 ± 0.1 14.3 ± 0.2 315 | Kanz 15 ^p | 42.0 ± 0.6 | 37.5 ± 0.6 | 511 | IIIc | -79 (39) |
| Kanz U9 ^α 41.8 ± 0.6 Kanz U9 ^T 42.2 ± 0.3 Kanz U10 ^P 43.7 ± 0.3 33.2 ± 0.5 2162 IIIc -73 (49) Kanz U11 37.9 ± 0.3 30.0 ± 0.3 1351 IIa -79 (27) Kanz U12 37.5 ± 0.7 31.7 ± 0.5 794 IIa -79 (26) Kanz U13 35.8 ± 0.4 33.8 ± 0.5 139 IIa -74 (26) Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II -71 (23) Kanz U16 ^P 30.4 ± 0.4 31.8 ± 0.3 -249 II -56 (29) Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II -80 (25) Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa -74 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc -68 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U29 37.6 ± 0.2 30.9 ± 0.4 31.8 ± 0.3 31.8 ± 0.3 IIIc -76 (30) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 30.9 ± 0.5 30.9 ± 0.4 30.9 ± 0.5 30.9 ± 0.4 30.9 ± 0.5 30.9 ± 0.5 30.9 ± 0.5 30.9 ± 0.4 30.9 ± 0.5 30 | Kanz U8 ^p | 37.0 ± 0.6 | 33.6 ± 0.4 | 333 | IIa | -77 (27) |
| Kanz U9 ^T 42.2 ± 0.3 Kanz U10 ^p 43.7 ± 0.3 33.2 ± 0.5 2162 IIIc -73 (49) Kanz U11 37.9 ± 0.3 30.0 ± 0.3 1351 IIa -79 (27) Kanz U12 37.5 ± 0.7 31.7 ± 0.5 794 IIa -79 (26) Kanz U13 35.8 ± 0.4 33.8 ± 0.5 139 IIa -74 (26) Kanz U14 ^T 34.8 ± 0.8 37.4 ± 0.4 -360 IIa -72 (26) Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II -71 (23) Kanz U16 ^p 30.4 ± 0.4 31.8 ± 0.3 -249 II -56 (29) Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II -80 (25) Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa -74 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc -68 (30) Bobo U21 33.4 ± 0.6 Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 33.2 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -77 (27) Others SEN(N)8 ^p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA GBR-A 17.0 ± 0.1 14.3 ± 0.2 315 | Kanz U9 | 42.3 ± 0.2 | 34.7 ± 0.5 | 1075 | IIa | -91 (28) |
| Kanz U10 ^p 43.7 ± 0.3 33.2 ± 0.5 2162 IIIc -73 (49) Kanz U11 37.9 ± 0.3 30.0 ± 0.3 1351 IIa -79 (27) Kanz U12 37.5 ± 0.7 31.7 ± 0.5 794 IIa -79 (26) Kanz U13 35.8 ± 0.4 33.8 ± 0.5 139 IIa -74 (26) Kanz U14 ^T 34.8 ± 0.8 37.4 ± 0.4 -360 IIa -72 (26) Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II -71 (23) Kanz U16 ^p 30.4 ± 0.4 31.8 ± 0.3 -249 II -56 (29) Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II -80 (25) Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa -74 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc -68 (30) Bobo U21 33.4 ± 0.6 Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U24 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U24 32.3 ± 0.5 24.4 ± 0.3 791 IIc -77 (27) Others SEN(N)8 ^p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA | Kanz U9α | 41.8 ± 0.6 | | | | |
| Kanz U11 37.9 ± 0.3 30.0 ± 0.3 1351 IIa $-79 (27)$ Kanz U12 37.5 ± 0.7 31.7 ± 0.5 794 IIa $-79 (26)$ Kanz U13 35.8 ± 0.4 33.8 ± 0.5 139 IIa $-74 (26)$ Kanz U14 ^T 34.8 ± 0.8 37.4 ± 0.4 -360 IIa $-72 (26)$ Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II $-71 (23)$ Kanz U16 ^p 30.4 ± 0.4 31.8 ± 0.3 -249 II $-56 (29)$ Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II $-80 (25)$ Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa $-74 (49)$ Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa $-76 (49)$ Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIc $-86 (20)$ Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIb $-103 (37)$ Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb $-84 (40)$ Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc $-68 (30)$ Bobo U21 33.4 ± 0.6 Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc $-76 (30)$ Bobo U24 32.2 ± 0.2 26.0 ± 0.3 791 IIc $-77 (27)$ Others SEN(N)8 ^p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA GBR-A 17.0 ± 0.1 14.3 ± 0.2 315 | Kanz U9 ^T | | | | | |
| Kanz U12 37.5 ± 0.7 31.7 ± 0.5 794 IIa -79 (26) Kanz U13 35.8 ± 0.4 33.8 ± 0.5 139 IIa -74 (26) Kanz U14 ^T 34.8 ± 0.8 37.4 ± 0.4 -360 IIa -72 (26) Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II -71 (23) Kanz U16 ^p 30.4 ± 0.4 31.8 ± 0.3 -249 II -56 (29) Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II -80 (25) Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa -74 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc -68 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 33.4 ± 0.6 Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 33.5 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U24 33.5 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U26 37.6 ± 0.2 26.0 ± 0.3 791 IIc -77 (27) Others SEN(N)8 ^p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA GBR-A 17.0 ± 0.1 14.3 ± 0.2 315 | Kanz U10 ^p | 43.7 ± 0.3 | 33.2 ± 0.5 | 2162 | IIIc | -73 (49) |
| Kanz U13 | Kanz U11 | 37.9 ± 0.3 | 30.0 ± 0.3 | 1351 | IIa | |
| Kanz U14 ^T 34.8 ± 0.8 37.4 ± 0.4 -360 IIIa -72 (26) Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II -71 (23) Kanz U16 ^p 30.4 ± 0.4 31.8 ± 0.3 -249 II -56 (29) Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II -80 (25) Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIIa -74 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIIc -68 (30) Bobo U21α 33.4 ± 0.6 Bobo U24α 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24α 33.0 ± 0.5 Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -76 (30) Bobo U20 31.5 ± 0.2 26.0 ± 0.3 791 IIIc -77 (27) Others SEN(N)8 ^p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA GBR-A 17.0 ± 0.1 14.3 ± 0.2 315 | Kanz U12 | 37.5 ± 0.7 | 31.7 ± 0.5 | 794 | IIa | -79 (26) |
| Kanz U15 33.4 ± 0.2 36.7 ± 0.3 -409 II -71 (23) Kanz U16 ^p 30.4 ± 0.4 31.8 ± 0.3 -249 II -56 (29) Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II -80 (25) Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIIa -74 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIIc -68 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 31.88 IIIc -77 (27) Others SEN(N)8 ^p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA GBR-A 17.0 ± 0.1 14.3 ± 0.2 315 | Kanz U13 | 35.8 ± 0.4 | 33.8 ± 0.5 | 139 | IIa | -74 (26) |
| Kanz U16 ^p 30.4 ± 0.4 31.8 ± 0.3 -249 II -56 (29) Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II -80 (25) Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIIa -74 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIIc -68 (30) Bobo U21 33.4 ± 0.6 Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 32.2 ± 0.2 27.6 ± 0.4 583 IIc -76 (30) Bobo U24 32.3 ± 0.5 Bobo U28 32.3 ± 0.5 24.4 ± 0.3 1388 IIc -85 (22) Bobo U30 31.5 ± 0.2 26.0 ± 0.3 791 IIc -77 (27) Others SEN(N)8 ^p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA GBR-A 17.0 ± 0.1 14.3 ± 0.2 315 | Kanz U14 ^T | 34.8 ± 0.8 | 37.4 ± 0.4 | -360 | IIa | -72 (26) |
| Kanz A 37.5 ± 0.2 28.5 ± 0.3 1696 II -80 (25) Bobongara Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa -74 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc -68 (30) Bobo U21α 33.4 ± 0.6 33.4 ± 0.6 33.4 ± 0.6 33.4 ± 0.6 33.0 ± 0.5 | Kanz U15 | 33.4 ± 0.2 | 36.7 ± 0.3 | -409 | II | -71 (23) |
| $\begin{array}{c ccccccccccccccccccccccccccccccccccc$ | Kanz U16 ^p | 30.4 ± 0.4 | 31.8 ± 0.3 | -249 | II | -56 (29) |
| Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa -74 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc -68 (30) Bobo U21 ^α 33.4 ± 0.6 33.4 ± 0.6 33.4 ± 0.6 33.4 ± 0.6 33.0 ± 0.5 | Kanz A | 37.5 ± 0.2 | 28.5 ± 0.3 | 1696 | II | -80 (25) |
| Bobo U10 37.2 ± 0.2 31.7 ± 0.4 731 IIa -74 (49) Bobo U11 37.9 ± 0.3 26.0 ± 0.3 2820 IIa -76 (49) Bobo U17 32.0 ± 0.2 26.6 ± 0.4 741 IIc -86 (20) Bobo U18 42.4 ± 0.2 32.7 ± 0.5 1897 IIb -103 (37) Bobo U20 37.6 ± 0.2 30.9 ± 0.4 1023 IIb -84 (40) Bobo U21 29.7 ± 0.2 26.3 ± 0.5 373 IIc -68 (30) Bobo U21 ^α 33.4 ± 0.6 33.4 ± 0.6 33.4 ± 0.6 33.4 ± 0.6 33.0 ± 0.5 | Bobongara | | | | | |
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| Others SEN(N)8p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA GBR-A 17.0 ± 0.1 14.3 ± 0.2 315 | Bobo U28 | 32.3 ± 0.5 | 24.4 ± 0.3 | 1388 | IIc | -85(22) |
| SEN(N)8p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA GBR-A 17.0 ± 0.1 14.3 ± 0.2 315 | Bobo U30 | 31.5 ± 0.2 | 26.0 ± 0.3 | 791 | IIc | |
| SEN(N)8p 52.4 ± 1.3 41.7 ± 0.8 2172 IIIc nA GBR-A 17.0 ± 0.1 14.3 ± 0.2 315 | Others | | | | | |
| GBR-A 17.0 ± 0.1 14.3 ± 0.2 315 | | 52.4 ± 1.3 | 41.7 ± 0.8 | 2172 | IIIc | nA |
| | ` ' | 17.0 ± 0.1 | 14.3 ± 0.2 | 315 | | |
| | GBR-A ^α | 17.0 ± 1.0 | | | | (-175) |

^aExcept where indicated, all samples are from the *Faviidae* family; the superscript p denotes *Porites* corals. Analyses on *Faviidae* corals are for the wall fractions only. The superscript α and T indicate previously reported values using α -counting and TIMS measurements, respectively (Chappell et al. 1996).

^bUsing Libby's half-life (5568 yr). The ¹⁴C-age is the weighted mean of ages from aliquots corresponding to the final 50% of coral mass selectively dissolved in at least three steps (Table 1).

^cThe Δ^{14} C value is the per mil difference between the initial 14 C/ 12 C of a sample (normalized for isotopic fractionation) and the 14 C/ 12 C of a standard; the standard has a value similar to the initial 14 C/ 12 C ration of 19th century wood. The Δ^{14} C is defined and discussed in (Stuiver and Polach 1977) and equals ($Fe^{\lambda T}-1$)×1000 per mil, where F is the fraction modern (Donahue et al. 1990), λ is the decay constant that corresponds to the 5730 year half-life, and T is the age in cal BP. In the present study, Δ^{14} C is calculated using the measured initial 14 C/ 12 C ratio and the 230 Th age. Quoted errors include errors in the measured 14 C/ 12 C ratio and 230 Th age.

^dReef numbering is described in (Ota et al. 1992; Chappell et al. 1996).

^eSea-levels obtained using both the height of corals as shown in brackets and uplift rates for both Kanzarua (2.8 m/ka) and Bobongara (3.3m/ka) reported by (Chappell et al. 1996).

dissolution. The weighted average is within twice the experimental error of individual measurements. Results from one coral from the Great Barrier Reef, originally dated by Veeh and Veevers (1970) with conventional ^{14}C and Th-U (13,600–13,860 \pm 0.220 ka BP and 17,000 \pm 1000 ka, respectively), are in good agreement with present results (14,390 \pm 0.150 ka BP and 16,981 \pm 0.112 ka). This sample is well within the field of other data of similar age (Figure 3). All of the other corals are from the uplifted Huon Peninsula coral terraces of Papua New Guinea collected during the 1992 international expedition (Ota et al. 1994). Four samples were previously Th-U dated using TIMS and five others by α -spectrometry (Chappell et al. 1996); they are included with the new results. The previous data have larger analytical uncertainties (Omura et al. 1995; Chappell et al. 1996), however, they are in good agreement with the new TIMS data.

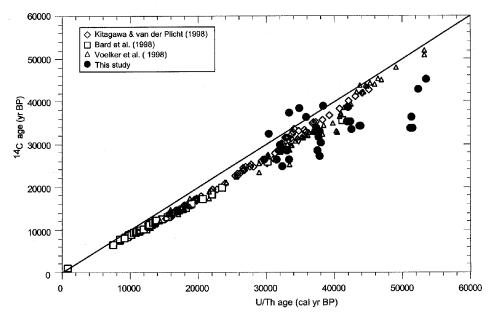


Figure 3 The present coral ¹⁴C and U-series ages, and previously published data. The baseline of younger ¹⁴C ages appear to be consistent in all the data sets. The present data show four distinct peaks of ¹⁴C excess. In between these peaks some ¹⁴C ages appear to be older than the corresponding calendar ages.

The measured Th-U ages of corals, from the uplifted Huon Peninsula Reefs II and III (Chappell 1974; Chappell et al. 1996), range between 29 ka and 55 ka. Several trends are apparent in the data; first, there is a noticeable increase in the divergence of the 14 C ages from calendar dates, from 30 ka to 50 ka, that follows the same trend as the bulk of the previous data; second, there are significant, 1000% to 7000% (equivalent to age shifts of 7 ka up to 17 ka) Δ^{14} C excursions at periods that correspond to particular reef-building episodes and sea-level rises at Huon Peninsula; and third, some 14 C ages in between the "peaks" are older than corresponding calendar ages.

That the Th-U ages cluster at particular times is no surprise as we are specifically dating corals from discrete reef sections; we do not have a continuous record of coral growth over this time period. However, the magnitude of the ¹⁴C age excursions from calendar ages is unprecedented. The present data are shown on an expanded scale in Figure 4. It is clear that data with ¹⁴C ages older than calendar ages

are distributed in-between the peaks. Two previous coral data points are within the range of present ages; one is a 30,200 yr old coral from Barbados, and the other a 42,000 yr old coral from Huon (Bard et al. 1993, 1998; Figure 3). Both of these overlap with the baseline defined by the present measurements (Figure 3). At tectonically uplifting sites, coral terraces grow when sea-level rise overtakes uplift. The uplift rate along the Huon coast is variable but typically in the order of 3 m per 1000 yr and the sea-level history has been derived previously (Chappell et al. 1996). The coincidence between sea-level rises and large excursions in ¹⁴C is, at first sight, puzzling. However, we believe that these events can be directly linked to rapid global climate changes, in particular, in the North Atlantic, and are also connected with Heinrich Events and periodic re-organizations of the North Atlantic thermohaline circulation.

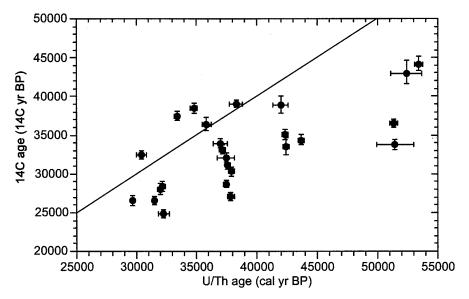


Figure 4 $\,^{14}C$ and Th-U ages plotted in an expanded scale diagram. Errors are 2σ mean. Four peaks of excess $\,^{14}C$ are clearly evident. There is no trend of increasing $\,^{14}C$ excess with age as the ≈ 37 ka peak is larger than the ≈ 43 ka peak. This, and other, detailed arguments based on tests of closed system behavior argue against an interpretation based on contamination by younger age or modern carbon as an explanation for the observed peaks.

There are large variations in the reported calendar ages of various Heinrich Events (Andrew 1998). This arises, in part, from mishandling of 14 C to calendar age conversions. Some 14 C ages are not even calibrated, but reported as absolute ages. Secondly, as the present data shows, the variability in 14 C ages may be due to rapid variations in atmospheric 14 C. We consider that data obtained by Bond et al. (1993) on the abundance of *Neogloboquadrina pachydermas* from the North Atlantic core V23-81 to represent one of the best examples of the timing and signature of Heinrich Events. The 14 C scale of this data set was converted to calendar ages using the prescription given by Bard et al. (1998). The results are shown in Figure 5 from a paper by Lund and Mix (1998). Inspection of Figure 5 shows that the 14 C peaks in the calibration curve match almost exactly with the timing of minima in the foraminiferal abundance and Heinrich Events. The 32 ka (calendar age) 14 C peak matches with H3. The 37.5 ka 14 C peak matches with the sharp cold stadial at 37.5 ka which is not marked as a Heinrich Event, but appears to belong to the same class. The 42.3 ka 14 C peak matches with H4 at \approx 42 ka, and the

52 ka ¹⁴C peak matches with H5 at 52 ka. This correspondence is remarkable and is highly unlikely to be coincidental.

The important point is that the Huon Peninsula terraces, which represent episodes of sea-level rise, appear to have been built as a direct consequence of large-scale discharges of icebergs into the North Atlantic and the associated sea-level rise. The present data assert that, concurrent with each episode, there was a huge rise in atmospheric 14 C. The 14 C excess, Δ^{14} C, is shown in Figure 5. The theoretical maximum 14 C excess in the atmosphere, if all the sinks are cut off, is 5870% (Siegenthaler et al. 1980) and 200% excess levels can be reached within a few hundred years. It appears that the present data provide direct evidence for the slow-down or shut-down of the North Atlantic thermohaline circulation as expounded by Broecker (1997, 1998). Furthermore, it solves a puzzle, again put forward by Broecker (1998):

Currently, ≈80% of the ¹⁴C atoms required to replace those undergoing decay in the deep sea enter via the Atlantic's conveyor circulation. Hence, if this route were to be shut down during the course of the 1200-calendar-year duration of the Younger Dryas, an enormous offset in the ¹⁴C timescale might be expected.

A few ¹⁴C ages, older than calendar ages, appear between the large ¹⁴C excursions. Presumably, they correspond to the resumption of the thermohaline circulation and transport of ¹⁴C to the deep ocean, initially at a rate greater than present. There is theoretical support for such age reversals in the calculations by Stocker and Wright (1998), when surface reservoir ages increase rapidly at the time of the commencement of the circulation. Interruption of the thermohaline circulation may be expected to redirect the northward movement of warm equatorial waters in the Atlantic to the Southern Ocean. Coupled with 5–10 m sea-level rise from North Atlantic icebergs, the Antarctic ice sheet may also be destabilized. The possibility that fresh water injection into Antarctic waters may interrupt the Antarctic deep-water formation has not been assessed in detail. However, Clark et al. (1996) consider a contribution to the melt-water-pulse-1A from the Antarctic ice sheet is most likely.

The 37 ka peak in the present data and the nominal 37 ka peaks in 10 Be and 36 Cl polar and deep-sea records, at first sight, appear to be generic. However, there is considerable uncertainty in the calendar age of the 10 Be/ 36 Cl peak. Various studies place this peak anywhere from 37 ka to 42 ka (Aldahan and Possnert 1998). It corresponds to a relatively long period when the geomagnetic field was low. Similarly the 10 Be peak is broad (≈ 3000 yr). Whether the 37 ka peak in the present data is the same as the 10 Be and 36 Cl peaks remains an open question.

Large Δ^{14} C peaks have been observed by Kitagawa and van der Plicht (1998) at 32–33 ka (Δ^{14} C up to 850%), by Voelker et al. (1998) at 33 ka (Δ^{14} C \approx 700%) and 39 ka (Δ^{14} C \approx 1200%), and by Bard (1998) at 41 ka (Δ^{14} C \approx 700%). Given the uncertainties in exact timing of these events, they could be related to the 14 C-excesses found in the present work. The magnitude of present Δ^{14} C range from 1000% to 2000% to 7000%. These values are within a factor of two to three of the previously determined highs. The difference between calendar and 14 C ages of the 51.5 ka peak is 17.8 ka. Due to the old calendar age of this and similar samples, they could have been highly susceptible to small amounts of contamination assuming the "true" 14 C ages were closer to 50 ka. However, a recent study by Schramm et al. (2000) of Lake Lisan sediments has identified a similar peak, U-Th dated at 51.2 ka with an age difference of greater then 9 ka, which tends to support the present findings.

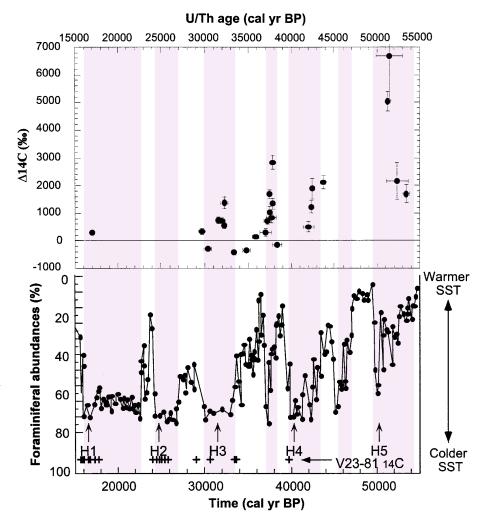


Figure 5 The upper panel shows the results from the present study. The lower panel is from the North Atlantic deep-sea core V23-81 showing systematic fluctuations in the abundance of *N. pachyderma* (s) correlated with the timing of Heinrich Events and cold stadials (Lund and Mix 1998). The time scale was derived from ¹⁴C data using the calibration recommended by Bard et al. (1993). Between H3 and H4, a brief duration sharp peak corresponds to an excess ¹⁴C peak found in the present data at 37.5 ka. This peak is not labeled as an Heinrich Event but appears to belong to the same class. Heinrich Events H3, H4, and H5 correlate with peaks of ¹⁴C. Such a correlation is unlikely to be coincidental and implies a genetic relationship between Heinrich Events, pulses of excess atmospheric ¹⁴C, sea-level rise and episodes of reef construction at Huon Peninsula. We surmise that the connecting factor between these events is the interruption of the North Atlantic Thermohaline circulation following periodic partial disruptions of the Laurentide ice-sheet.

CONCLUSION

More than 250 AMS ¹⁴C dates and 25 high-precision TIMS U-series ages have been determined for 30,000–50,000-yr-old corals from the uplifted terraces of Huon Peninsula in Papua New Guinea. Previously extensive efforts to define a ¹⁴C calibration curve for the period beyond around 12,000 yr, through directly measured calendar ages, only yielded a few scattered points. A new CO₂ extraction and graphatization line was constructed and used for processing "old" carbonate samples only. ¹⁴C

dates of samples older than 20,000 yr may be affected by secondary younger carbon contamination. For this reason, each coral sample was subjected to partial dissolution, and the results were considered acceptable only when the 14 C ages of at least three aliquots were self consistent within errors. This is the only quantitative test of diagenetic alteration in corals for 14 C dating purposes. In addition, conventional screening methods were applied, including textural investigation of thin sections under a petrographic microscope, XRD, and stable isotope analyses. For these reasons, results obtained in the present study are considered as reliable as can be ascertained with presently available methods. However, in some circumstances, it is still possible that contaminated samples could pass through the above tests. For example, the high 14 C excesses found in ≈ 50 ka (calendar age) corals, although supported by recent dating of Lake Lisan sediments (Schramm et al. 2000), may need further work for confirmation. In addition to tests for 14 C contamination, the samples were tested for open system behavior for U and Th isotopes, such as total U and 232 Th concentrations, and δ^{234} U(T) values. Most samples were within acceptable limits of these tests.

The results of the 14 C time-scale calibration experiments can be summarized as follows: First, the trend for increasing divergence between 14 C ages and calendar ages appears to continue beyond 30 ka and up to 50 ka; the limit of present measurements. At the younger end, data from the present study overlap with previous data. Second, four discrete age peaks of enhanced 14 C concentration were observed. The clustering in age is expected as the samples come from distinct, individual reef sections. However, the prominent peaks of Δ^{14} C, including some negative values, in between the peaks, were unexpected. It is possible that the peaks are due to secondary, younger age, or modern carbon contamination. While this is acknowledged, and some of the older \approx 50 ka samples may have been affected, it is unlikely to be the dominant cause given the battery of tests the samples were required to pass. The negative values of Δ^{14} C, that systematically cluster in between the peaks, could not be produced by such contamination.

The four ¹⁴C peaks appear correlated with the timing of North Atlantic Heinrich Events and ice-rafted debris deposits in deep sea cores. The accepted explanation for these Dangaard-Oeschger and Bond cycles involve reorganizations in the global ocean circulation system. The North Atlantic deep-water formation during glacial periods appears to be sensitive to fresh meltwater input into the North Atlantic by periodic instabilities in the ice sheets that launched armadas of icebergs into Hudson Bay and resulted in sea-level rises. Slow-down or closure of the circulation is expected to significantly raise atmospheric ¹⁴C levels. The present data show an initial high sea-level followed by steep rise in atmospheric ¹⁴C levels as the sea-level falls. The high concentration lasts for less than ≈1000 yr and drops below original ambient levels over several thousand years. The negative values of Δ^{14} C indicate the vigorous restart of the NADW circulation or enhanced ventilation at other oceans. Deep water formation around the Antarctic continent is believed to be equal to that in the North Atlantic. Cessation of the North Atlantic circulation is likely to redirect warm equatorial waters to the Southern Ocean and coupled with a high sea-level pulse is likely to destabilize the Antarctic ice sheet as well. 14C and Th-U results obtained in the present work constitute a direct confirmation of severe and abrupt climate oscillations observed in Greenland ice cores and deep sea sediments. They provide direct evidence for the start-stop behavior of the North Atlantic conveyor belt circulation during the last glacial as put forward mainly by Wallace S Broecker (1998), and the increasingly evident global influence of these events beyond the boundaries of North Atlantic.

ACKNOWLEDGMENTS

We thank GS Burr, PU Clark, and J Jouzel for their helpful comments, and J van der Plicht for editorial assistance. Some of the coral samples were AMS dated at ANSTO supported by AINSE grant (97/060R).

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