

MODELLING SEA ICE GROWTH

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Abstract

The freezing of water to ice is a classic problem in applied mathematics, involving the solution of a diffusion equation with a moving boundary. However, when the water is salty, the transport of salt rejected by ice introduces some interesting twists to the tale. A number of analytic models for the freezing of water are briefly reviewed, ranging from the famous work by Neumann and Stefan in the 1800s, to the mushy zone models coming out of Cambridge and Oxford since the 1980s. The successes and limitations of these models, and remaining modelling issues, are considered in the case of freezing sea-water in the Arctic and Antarctic Oceans. A new, simple model which includes turbulent transport of heat and salt between ice and ocean is introduced and solved analytically, in two different cases—one where turbulence is given by a constant friction velocity, and the other where turbulence is buoyancy-driven and hence depends on ice thickness. Salt is found to play an important role, lowering interface temperatures, increasing oceanic heat flux, and slowing ice growth.

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1. Introduction

The temperature at which sea-water freezes depends on its salinity. The saltier the brine, the lower the freezing point. When salinity is above the critical value of 24 psu (practical salinity units), the freezing point is below the temperature of maximum density, so that colder sea-water is heavier. Then cooler surface waters are heavier and overturn, mixing a region of ocean down to a critical depth (the pycnocline). This region is cooled to near freezing, rather than just the surface waters.

When cold air causes open sea-water to begin to freeze, many small ice crystals (frazil ice) form at the surface, stirred by wind, waves and currents [6, 10, 13]. This may be further complicated by the formation of agglomerations of crystals called

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pancake ice, which collide and raft. A large solid sheet of ice is typically formed when a critical thickness of about 10 cm is reached [6]. Then those crystals that are oriented with the direction of most rapid growth near vertical become predominant as the ice thickens.

Salt is rejected by growing ice as a dense brine, but the brine cannot escape as rapidly as heat at the ice–ocean interface. This is due to the diffusivity of salt being much less than thermal diffusivity. Then the very cold salty interface leads to a cold fresher region nearby. As a consequence, while the actual ice–ocean interface is at freezing point, the region below the interface is below freezing point, creating an unstable situation—any ice that protrudes into the supercooled region grows faster. This leads to the dendritic growth of ice as fingers or plates down into the sea. These fingers and plates then bridge across, trapping brine in the ice–brine mixture. This constitutional supercooling causes the ice–ocean interface to be convoluted, with a gradual transition from water to a mixture of liquid (brine) and ice called a mushy zone [7, 10, 34, 35], rather than a sudden planar change from liquid to ice.

Cold air drives the freezing process, and snow on top can complicate it by decreasing conductive heat flow. There is radiative heat transfer between atmosphere and ice. The ocean is relatively warm, with turbulent currents, and in Antarctica near glacial ice shelves is occasionally slightly supercooled with billows of small ice crystals present, to further complicate the picture.

2. Previous models

Classic work by Neumann in the 1860s [3] and, famously, by Stefan in 1891 [13] was seminal in the early modelling of the growth of sea ice. The diffusion equation, resulting from conservation of heat, is solved in one dimension with a moving boundary between ice and ocean, in the case where the boundary is sharp and heat transport from ocean to ice is negligible. The freezing interface is found to move as the square root of time. Salt transport is ignored.

Conservation of energy leads to the heat conduction equation for the temperature $T(t, z)$ of the ice ($^{\circ}\text{C}$), assuming that the ice can be approximated as a one-dimensional sheet with planar interfaces with air and ocean, z is elevation and t is time:

$$\frac{\partial T}{\partial t} = D \frac{\partial^2 T}{\partial z^2}, \quad D \equiv \frac{k_i}{\rho_i C_i}. \quad (2.1)$$

Here D is the thermal diffusivity of the ice (which averages about $10^{-6} \text{ m}^2 \text{ s}^{-1}$ over the temperature range -2 to -25°C for an ice salinity of 5 psu), $\rho_i \approx 910 \text{ kg m}^{-3}$ is the ice density, C_i its thermal capacity ($\text{J kg}^{-1} \text{ }^{\circ}\text{C}^{-1}$), and $k_i \approx 2.2 \text{ W m}^{-1} \text{ K}^{-1}$ its thermal conductivity [18, 24, 25]. Trapped brine means that C_i and D have a strong dependence on T [23, 24].

Boundary conditions used by Stefan and Neumann are that at the air–ice interface $z = 0$ the ice temperature T_0 is equal to the air temperature $T_a(t)$, and at the moving boundary $z = -h(t)$ that is the ice–ocean interface, the temperature is at freezing

point T_f , and the latent heat removal that advances the freezing front at \dot{h} m s⁻¹ is caused by heat conduction upwards to the colder air through the ice, so that

$$\rho_i L \dot{h} = -k_i \left. \frac{\partial T}{\partial z} \right|_{-h^+}$$

where $L \approx 3.3 \times 10^5$ J kg⁻¹ K⁻¹ is the latent heat of fusion for sea-water.

The air temperature oscillates, and these temperature variations penetrate sea ice at rates (and to effective depths) dependent on frequency, with a phase velocity [3] estimated for constant diffusivity as $2D/\delta$ m s⁻¹ where δ is the penetration depth of the thermal wave. The slowest waves then penetrate with phase velocities $2D/h$ m s⁻¹, which is faster than the ice growth rate estimated at $k_i(T_f - T_a)/(\rho_i L h)$ provided that

$$T_f - T_a < \frac{2D\rho_i L}{k_i} \approx 300^\circ\text{C}.$$

For air temperatures encountered in Antarctica and the Arctic, this inequality is easily satisfied. T_f is about -2°C depending on salinity, and T_a during freezing ranges from this value to about -70°C . This means that the slowest penetrating thermal waves do travel much faster than the ice growth rate. This does not mean that the temperature profile is always linear—observations indicate otherwise, and the observed curvature of temperature versus depth has been used to directly estimate the thermal conductivity of sea ice [18, 24, 25]. However, oscillations also damp out as they penetrate, with a damping factor of the order of $\exp(-\sqrt{\omega/(2D)}z)$ [3], where ω is the frequency of oscillation. Then, for example, an oscillation with period 40 days damps by a factor of e over a lengthscale of 1 m, and smaller period oscillations damp out over even shorter depths. This suggests that a running average of 40 days or so is a reasonable pre-treatment for air temperatures—the ice effectively averages out rapid changes. Such a running average has much less variability than the original air temperature.

For the purposes of calculating ice growth rate, a linear approximation (that is, a steady-state solution to (2.1)) is often used to estimate the average temperature gradient in the ice at the ice–ocean interface,

$$\frac{\partial T}{\partial z} \approx \frac{T_a - T_f}{h},$$

and the Stefan problem becomes

$$\dot{h} = \frac{k_i}{\rho_i L h} (T_f - T_a).$$

This integrates to give the Stefan solution

$$h^2 = \frac{2k_i}{\rho_i L} \theta \approx 1.4 \times 10^{-8} \theta,$$

where $\theta \equiv \int_0^t (T_f - T_a) dt$ is the number of degree-seconds of cooling. For a constant air temperature θ is linear in time, so that h grows as the square root of time. The Stefan solution is graphed in Figure 2, alongside data and later model results.

Stefan's solution is generally regarded as giving an upper limit on ice growth rate and thickness. Snow reduces heat flow, solar heating slows cooling, and heat flow from ocean to ice slows cooling—these effects have been ignored.

A more sophisticated boundary condition at the upper ice surface allows for a thermal boundary layer in the air, or for a layer of snow, so that there is a flux (or mixed boundary) condition,

$$\frac{\partial T}{\partial z} = \alpha(T_a - T_0), \quad z = 0,$$

where α is a nonnegative heat transfer coefficient with units per metre, which depends on snow thickness if snow is present [12, 16]. This results in

$$\dot{h} = \frac{k_i \alpha}{\rho_i L(1 + \alpha h)}(T_f - T_a),$$

which integrates to give

$$h^2 + \frac{2}{\alpha}h = \frac{2k_i}{\rho_i L}\theta. \quad (2.2)$$

In the limit $\alpha \rightarrow \infty$, the mixed boundary condition reduces to the boundary condition used by Stefan, $T_a = T_0$, and the solution for h reduces to the Stefan solution. In general for positive α , ice growth is slower than the square-root-of-time Stefan behaviour, although for $\alpha = 40 \text{ m}^{-1}$ the graphs practically coincide (Figure 2). Small α gives linear growth of h with θ , and zero α corresponds to a perfectly insulated upper boundary, with no heat taken from the ice and no ice generated.

An empirical fit [1] to the thickness of snow-free ice in Thule, Greenland, illustrated in Figure 1 to show the scatter of data points, used the same formula,

$$h^2 + 0.051h = 0.775 \times 10^{-8} \theta \quad (2.3)$$

where h is in metres and θ is degree-seconds. See [17] for other empirical relationships.

Matching the coefficients of h in (2.2) and (2.3) implies that $\alpha \approx 40 \text{ m}^{-1}$. However, the constant term on the right-hand side of (2.3) is not a good match with known values listed earlier in this paper. Figure 2 shows a comparison between the empirical (data) thicknesses given by (2.3) and our solution (2.2), using $\alpha = 40 \text{ m}^{-1}$. Our simple model (with up-to-date values for k_i , ρ_i , and L) predicts a faster growth of sea ice than is observed.

Reducing our k_i value from 2.2 to 1.3 gives a good fit to the data (graphically indistinguishable from the fit that Anderson drew; Figure 1), but this is at the cost of almost halving the typical measured value of thermal conductivity in sea ice [18]. A number of missing factors in the model could account for the misfit, including radiative heat transfer, heat flow from the ocean to the ice (the oceanic heat flux), and the role played by the transport of salt. We will consider the latter two mechanisms in the next section.

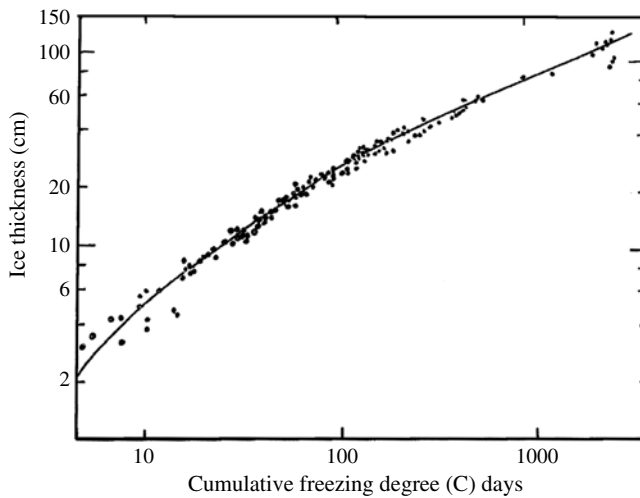


FIGURE 1. Data and empirical fit to snow-free ice in Greenland, after Anderson, 1961 [1]. (Reprinted with permission of the International Glaciological Society and Prof. Anderson.)

Also shown in Figure 2 is a fit to sea ice thicknesses measured in McMurdo Sound [26]. These grow initially more slowly than Anderson's measurements, then more rapidly once the ice is about 1 m thick. This may be related to the appearance of a different ice crystal structure in the McMurdo ice [26] at this thickness, possibly associated with the accretion of ice from a supercooled mixed oceanic layer. Slow early solid ice growth may be due to heat transport from this layer, as it is supercooled, and later faster growth of solid ice may be reaping the benefit of small crystals of ice that are nucleated in the mixed layer, rising to the solid ice interface and sticking there.

A very detailed vertical model of the heat conduction problem is solved numerically by Maykut and Untersteiner [17]. They concentrate particularly on the atmosphere–ice interface, and on parameterizations of the various mechanisms for heat exchange there, especially radiative transport. Salt transport is not modelled, however, which may impact on heat transport from the ocean.

3. Salt transport

Here we consider the importance of also modelling the transport of salt, for the growth of sea ice. Salt alters the freezing point of sea-water, and salt diffusivity is very different from thermal diffusivity.

Weber [32] considers the very early growth of sea ice, and includes salt transport. He shows that the coupling between salinity and heat transport cannot be neglected. His model, like that of Stefan, ignores heat transport between ice and ocean. Salt transport to the ocean is modelled as diffusive. The model is solved by a heat balance integral method, which uses a polynomial approximation to the temperature and salinity profiles. Weber assumes a small temperature drop across the ice, and takes

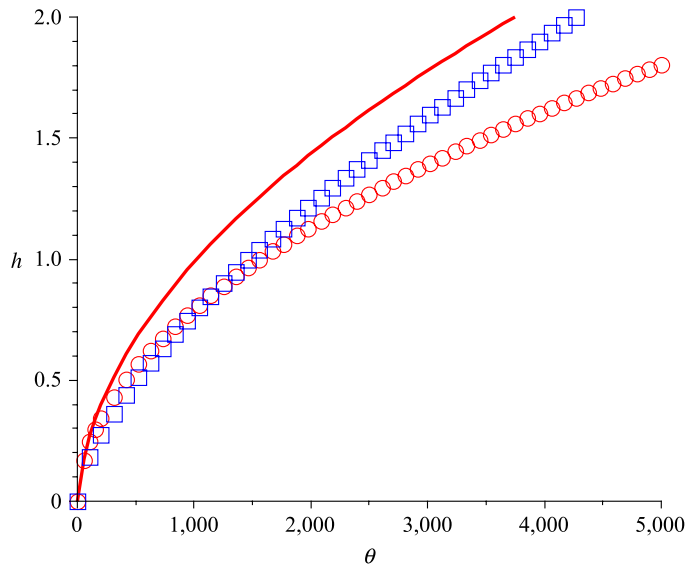


FIGURE 2. A graph of ice thickness h in metres versus degree-days of cooling θ , comparing an empirical fit by Anderson [1] (circles), an empirical fit to data from McMurdo Sound [26] (boxes), and Stefan's solution and our simple model (2.2) with $\alpha = 40 \text{ m}^{-1}$, which both plot almost identically (solid line).

the interface to be sharp and smooth. Recent measurements and modelling suggest that all three of these assumptions are violated in practice (for example [6]) in the early growth of fresh polar sea ice.

Weber does find that including salt transport reduces ice growth rate, because salt lowers the freezing temperature at the ice–ocean interface, and for his thin ice this affects heat transport to the atmosphere. Furthermore, ice growth rate in his model is proportional to the square root of time, for constant air temperature. Weber notes that the rejected salt will lead to convection, and that morphological instability and dendritic ice growth are likely, invalidating the assumption of a flat interface between ice and ocean.

A similar approach is taken by Notz *et al.* [22] in their modelling of the formation of a false bottom under summer melt ponds in the Arctic. Their model admits a similarity solution, since there are no external length or timescales, and interfaces grow as the square root of time.

3.1. Turbulent transport The turbulent transport of heat and salt in the general context of cooling a binary alloy is modelled by Woods and Huppert [33]. They posit a boundary layer beneath the ice, and flux terms depending on a Rayleigh number to the power of one third. Various turbulent sub-models are considered, including the formation of plumes or blobs of brine, and various relative rates of transport of heat and salt are examined. Their mixed layer is of finite extent, and may become supercooled. Small-time asymptotic analysis gives square-root-of-time growth, which changes to

linear-in-time behaviour later on, in the limit of relatively slow salt transport.

A new simple model that extends the Stefan approach to include the turbulent transport of heat and salt at the ice–ocean interface is now introduced here and solved approximately. A linear temperature profile in the ice is assumed, with the surface of the ice at air temperature. Conservation of heat and salt at the growing interface gives

$$\dot{h} = \frac{k_i(T_f - T_a)}{\rho_i L h} - \frac{Q_T}{\rho_i L} \quad (3.1)$$

and

$$\dot{h} = \frac{Q_S}{\mathcal{D}S}, \quad (3.2)$$

where $\mathcal{D}S = \rho_i(S_i - S_w)$, and Q_T and Q_S are turbulent fluxes of heat and salt.

We assume that there is a thermal and saline boundary layer in the ocean next to the ice with density ρ_w , temperature T_f , and salinity S_w . The average salinity in the sea ice is S_i . Units for salinity are parts per thousand, or grams of salt per kilogram of brine. We assume that parts per thousand equates to psu.

The average salinity of sea ice and of the ocean, near the ice–ocean interface, depend on how fast it is freezing. We model the variation of sea ice salinity by taking the salinity of sea ice to be proportional to the (variable) salinity of the sea-water near the ice (as in [27]), so that

$$S_i = f S_w \quad \text{where } f \approx 0.14,$$

and hence,

$$\mathcal{D}S = -S_w \varrho_i (1 - f) \approx -0.86 S_w \varrho_i.$$

Freezing point is approximated as [32]

$$T_f = -A S_w,$$

where $A \approx 0.054^\circ\text{C}/\text{psu}$.

The turbulent flux terms for heat and salt are (using a simple model for transfer from a region close to the ice with salinity and density S_w and ϱ_w , to a well-mixed region further away)

$$Q_T = -C_H u_* (T_f - T_m) \varrho_w C_w \quad \text{and} \quad Q_S = -C_S u_* (S_w - S_m) \varrho_m,$$

where the constants T_m , ϱ_m , and S_m are temperature, density, and salinity in a deeper, well-mixed ocean layer, C_w is the thermal capacity of that water, and $C_H \approx 0.0058$ is a turbulent transfer coefficient measured in the field [19]. A value for C_S is here calculated from C_H in the spirit of [20] as

$$C_S = C_H \left(\frac{\text{Pr}}{\text{Sc}} \right)^{2/3} \approx 0.03 C_H \approx 2 \times 10^{-4}.$$

Here, $Sc = 2432$ is the Schmidt number for salt in water (kinematic viscosity divided by molecular diffusivity), and $Pr = 13.4$ is the Prandtl number (kinematic viscosity divided by thermal diffusivity). This parameterization of the turbulent transport of salt and heat reflects the importance under ice of a viscous sublayer, across which molecular properties are significant [20], making the transport of salt comparatively slow.

We will use the simple average value for the friction velocity, $u_* \approx 0.015 \text{ m s}^{-1}$, as measured in the Weddell Sea in 1996 [19], and take the mixed layer in the ocean to be at the freezing temperature for its salinity. Then our variables are ice thickness $h(t)$ and boundary-layer salinity $S_w(t)$. Equating the right-hand sides of (3.1) and (3.2) gives a quadratic for S_w as a function of h ,

$$S_a S_w^2 + S_b S_w + S_c = 0, \quad (3.3)$$

where

$$S_a \equiv \frac{C_H u_* \rho_w C_w A}{L} + \frac{k_i A}{Lh}, \quad S_b \equiv \frac{C_H u_* \rho_w C_w T_m}{L} + \frac{k_i T_a}{Lh} + \frac{C_s u_* \rho_m}{0.86}$$

and

$$S_c \equiv \frac{-C_s u_* \rho_m S_m}{0.86}.$$

With an eye on (3.2), it is useful to transform variables in (3.3) as $Y = S_w / (S_w - S_m)$. The resulting quadratic in Y ,

$$Y_a Y^2 + Y_b Y + S_c = 0,$$

where $Y_a \equiv S_a S_m^2 + S_b S_m + S_c$ and $Y_b \equiv -(S_b S_m + 2S_c)$, has a solution that is closely matched by taking the balance between the Y^2 and the Y terms, giving a linear dependence on h ,

$$Y \approx \frac{T_a}{AS_m + T_a} + \frac{u_*}{k_i} \left(\frac{C_H \rho_w C_w T_m - (C_s \rho_m L / 0.86)}{AS_m + T_a} \right) h \approx 1 + Y_1 h,$$

where $Y_1 \approx -780/T_a$ for air temperatures much less than T_f .

Then if the air temperature T_a is taken to be constant, (3.2) integrates to give the solution

$$h^2 + \gamma_1 h \approx \frac{2k_i}{\rho_i \mathcal{L}} \theta,$$

where $\gamma_1 \equiv 1.7k_i |T_a| / (u_* C_s \rho_m \mathcal{L})$ and $\mathcal{L} \equiv L - (0.86 C_H \rho_w C_w T_m) / C_s \rho_m$. Using our parameter values this becomes

$$h^2 + \frac{|T_a|}{460} h \approx 0.88 \times 10^{-8} \theta, \quad (3.4)$$

which, using a range of air temperatures, $T_a = -5$ to -45°C , gives a close match to Anderson's data as illustrated in Figure 3.

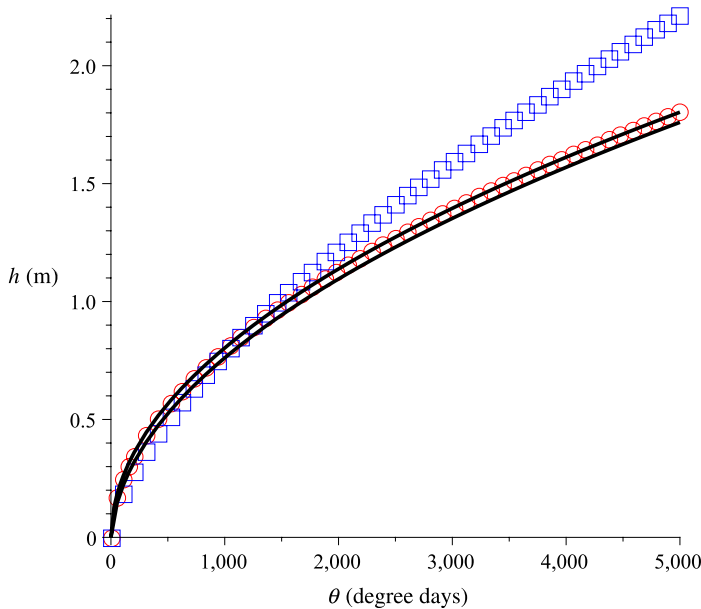


FIGURE 3. Data from Anderson [1] (circles), Purdie *et al.* [26] (boxes), and our turbulent salt transport model (3.4) with $T_a = -5^\circ\text{C}$ (upper curve) and $T_a = -45^\circ\text{C}$ (lower curve).

The term γ_1 reflects the effect of salt transport on ice growth, and vanishes as $C_s u_* \rightarrow \infty$ and salt transport increases. If $C_s \rightarrow \infty$, then $\mathcal{L} \rightarrow L$ and Stefan's solution is recovered. If C_s vanishes, h becomes linear in θ and also vanishes.

Note that salt transport has considerably reduced the modelled ice growth rate. Other authors have noted the importance of including salt transport, although in different contexts [22, 32] to the present one. Here, salt can be understood to be a bottleneck in the transport process, accumulating in our model at the ice–ocean interface because it has a much smaller diffusivity than heat. The heat balance is satisfied because the accumulation of salt lowers the ice temperature there, and increases the oceanic heat flux to the ice (commensurate with a slower growth rate).

The heat flux from the freezing interface to the air, and the heat flux from the ocean to the freezing interface, are graphed in Figure 4, assuming an air temperature of -25°C . These values compare well with measured values in polar waters [19]. The oceanic heat flux is significant compared with the heat flux to the atmosphere, which is about 2.6 times the oceanic heat flux.

The ratio of oceanic and atmospheric heat fluxes is almost constant over the ice thickness range 0.1–2 m. This is because both fluxes depend approximately on the inverse of ice thickness h , the flux to air because the average temperature gradient is inversely proportional to h , and the oceanic heat flux because the salinity and hence the freezing temperature at the interface is roughly inversely proportional to h .

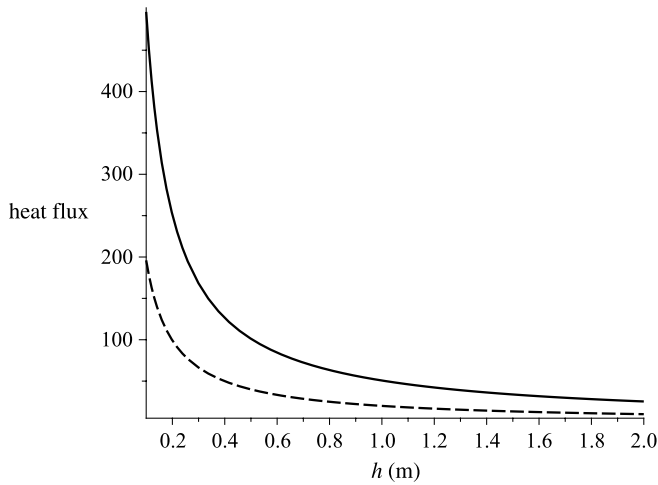


FIGURE 4. Heat fluxes (W m^{-2}) from freezing ice–ocean interface to air (upper curve) and from the ocean to that interface (lower curve), versus ice thickness, for the turbulent transport model with constant friction velocity and an air temperature of -25°C . The net heat flux from the freezing front is the difference between these two fluxes.

3.2. Buoyancy-driven turbulence The previous model used a constant friction velocity that reflected an average turbulence environment consistent with field measurements. However, one of the factors driving turbulent transport is the rejection of brine from growing sea ice in turbulent plumes. Since brine is rejected more rapidly when ice is growing faster, there is a coupling between friction velocity and ice growth rate. We now consider the effects of taking turbulent transport to be driven entirely by brine rejection.

This may be modelled [2, 8, 14, 31] by taking $u_* = \sqrt{C_d} W_*$, where $C_d \approx 1.3 \times 10^{-3}$ is the drag coefficient [9], and W_* is the Deardorff or free convection velocity scale [5]. Turner [30] gives W_* in terms of the buoyancy flux

$$W_*^3 = B\lambda,$$

where the buoyancy flux is $B \equiv -g \langle \rho'_w w' \rangle / \langle \rho_m \rangle$ (see also [21]), $\langle \rho'_w w' \rangle$ is the average vertical flux (positive upwards in this paper) of random variations ρ'_w in fluid density at the ice–ocean interface, and $\langle \rho_m \rangle$ is the mixed layer density. The lengthscale used is [21] $\lambda = \kappa z_i$ where $\kappa = 0.4$ is von Kármán’s constant. The term z_i is the thickness of the mixed layer, stirred by brine rejection under the growing sea ice, and from measurements [11] is of the order of 50 m in mid-winter in McMurdo Sound. Fluid density depends mainly on salinity, so that

$$\frac{d\rho_w}{dS} \equiv \rho_{ws} \approx 0.81 \text{ kg/m}^3/\text{psu}$$

and $\rho'_w = \rho_{ws} S'_w$. Then $B \approx -\rho_{ws} g \langle w' S'_w \rangle / \langle \rho_m \rangle$. The average vertical flux of salt

follows from our model, using (3.2):

$$\langle w' S'_w \rangle = \dot{h} \frac{DS}{\rho_w} \approx - \frac{0.86 \dot{h} S_w \rho_i}{\rho_w}.$$

Hence,

$$W_*^3 = 0.86 \rho_{ws} \lambda g \dot{h} \left(\frac{S_w \rho_i}{\rho_m \rho_w} \right).$$

Then the friction velocity may be found in terms of salinity in the boundary layer from the model equations, $u_* = C_d^{3/4} \sqrt{b(S_w - S_m)}$ and $W_* = u_* / \sqrt{C_d}$, where $b \equiv \lambda g C_S \rho_{ws} / (0.86 \rho_w)$.

Equating the right-hand sides of (3.1) and (3.2) then gives a quintic in $\Delta S \equiv \sqrt{S_w - S_m}$. The physically realistic solution to this quintic is approximated within 10% by matching the cubic and constant terms, to get a linear relationship between ΔS and $h^{-1/3}$. Substituting this back into (3.2) gives

$$\dot{h} \approx \frac{k_i |T_a|}{\rho_i L \gamma_3 (h + \gamma_2 h^{1/3})},$$

where

$$\gamma_2 \equiv \left(\frac{0.86 k_i |T_a|}{\gamma_3 C_s \rho_m L C_d^{3/4} \sqrt{b S_m}} \right)^{2/3} \quad \text{and} \quad \gamma_3 \equiv 1 - \frac{0.86 C_H \rho_w C_w T_m}{C_S \rho_m L}.$$

This integrates for constant T_a , giving

$$h^2 + \frac{3}{2} \gamma_2 h^{4/3} = \frac{2k_i}{\rho_i L \gamma_3} \theta. \quad (3.5)$$

The term γ_2 is of order C_s^{-1} , so vanishes as salt transport increases, recovering Stefan's solution, and γ_2 is singular as salt transport reduces to zero, giving similar behaviour to the previous model which assumed constant u_* . The solution given by (3.5) is graphed in Figure 5 for the case where air temperature is -25°C , and compared with data and the previous model.

The modelled ice thickness when turbulence is driven by buoyancy is much less than data and other models indicate. This is because of the assumption that only buoyancy, from rejected brine plumes, drives turbulence under sea ice. Typical values of u_* obtained from this model are about 100 times smaller than the measured value $1.5 \times 10^{-2} \text{ m s}^{-1}$ used in the previous model. This suggests that brine rejection plays a minor role in generating turbulence in the mixed layer. So in this model, salt builds up at the ice–ocean interface, lowering the temperature there. The result is a much increased oceanic heat flux to the freezing interface, and hence a much reduced net heat flux from it. Relatively more heat is removed from the oceanic mixed layer, at the expense of the growth of the freezing front.

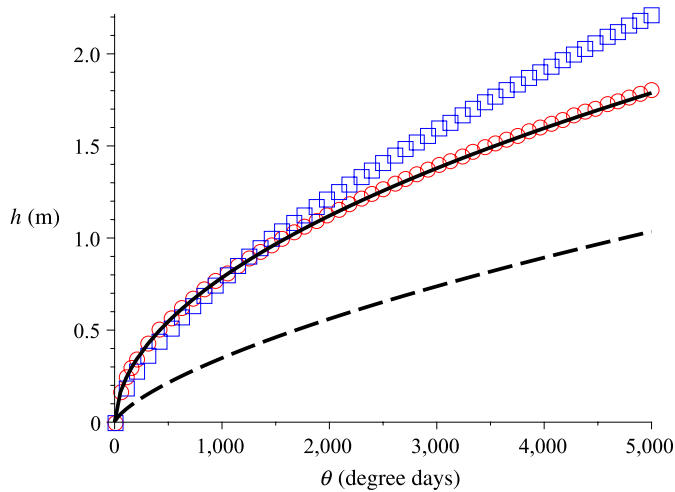


FIGURE 5. Data from Anderson [1] (circles), Purdie *et al.* [26] (boxes), a constant turbulence salt transport model (3.4) with $T_a = -25^\circ\text{C}$ (solid curve) and a model with only buoyant plumes driving the turbulence (dashed curve) as in (3.5), also with $T_a = -25^\circ\text{C}$.

Both turbulent salt flux models predict cooling of the oceanic mixed layer by this mechanism, which may lead to supercooling. It is unclear how supercooling in the mixed layer is relieved—ice may grow in dendritic structures from the interface, especially if supercooling is greater nearby, and/or ice may grow from nucleation sites within the turbulent mixed layer. The latter mechanism is a relatively efficient way to make ice, in the form of turbulent billows of small ice crystals [11, 15] within the mixed layer, as this ice does not immediately thicken the layer of solid ice. Hence heat flow to the atmosphere is more rapid than if the ice does thicken this layer.

A dendritic response to supercooling means the interface between ice and ocean is not planar but highly convoluted. One modelling approach to this phenomenon is to approximate such a mushy zone as a porous medium made of ice and brine, with a porosity and permeability that depend on temperature [7, 33–36]. The ice–ocean boundary is then warmer than in the planar interface models with salt transport, and oceanic heat flux is reduced.

4. Conclusions

The inclusion of salt transport is important when modelling the growth of sea ice. A simple model for turbulent transport of salt and heat between ice and ocean characterized by a constant friction velocity has been developed and solved. In this model, the presence of salt at the ice–ocean interface lowers the freezing temperature there, increasing the heat flux from the ocean, and slowing down the growth of sea ice, to give an excellent match with observations.

While the modelling of the growth of sea ice has come a long way since the remarkable work of Neumann and Stefan, it still falls short of modern understandings of ice structure and oceanic conditions. The constant friction velocity model presented here does well in fitting Anderson's classic data, and is consistent with observed oceanic heat fluxes, but fails to explain changes in ice crystal structure often seen in thicker sea ice [28, 29].

Mushy zone models are a good approach to modelling the dendritic growth observed in supercooled waters, but do not address the question of which ice growth mechanism will dominate in a supercooled ocean under sea ice—dendritic growth from solid surface ice, or frazil growth from nucleation sites within a turbulent oceanic mixed layer. Models addressing such frazil growth will have to contend with the growth and breakage of a population of small discs of ice with a size distribution [4], rising under buoyancy but stirred by turbulence, stabilizing the turbulence by buoyancy, and interacting with the temperature and turbulence fields and with the solid ice above.

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